K ink m anifolds in a three-com ponent scalar eld theory

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A bstract

In this work we identify the manifold of solitary waves arising in a three-component scalar eld model using the Bogom ol'nyi arrangement of the energy functional. A rich structure of topological and non-topological kinks exists in the dierent sub-models contained in the theory.

1 Introduction

The search for solitary waves is an ongoing topic in both M athematics and Physics because this kind of quasi-soliton plays an important rôle in a huge number of branches of non-linear science. In Field Theory, they usually appear in models that support spontaneous sym m etry breaking, the most prominent examples being kinks/domain walls, vortices, and m onopoles [1]. Starting with theories that involve a high number of elds, the usual procedure followed to investigate the existence of solitary waves -topological defects- is to obtain an e ective scalar eld theory, in posing severe restrictions on the original theory. In most cases, one is compelled to pursue an elective theory that will correspond to a single scalar eld model, where the existence of topological defects can be checked easily. The reason for this is the good-understanding of this kind of system; as paradigm atic examples, the well-known kink and soliton in the one-dimensional and sine-Gordon models should be noted. Both kinds of solitary wall can be thought of as thick walls, the topological defects in a three-dimensional perspective. Nevertheless, the general fram ework is that the elective theory depends on several scalar elds and thus the truncation may involve an important loss of information concerning the presence of topological defects and the structure of spontaneous symmetry breaking. It is therefore desirable to investigate the general properties of dom ain wall solutions in a multi-scalar eld theory.

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In (1+1)-dimensional eld theory, solitary waves are non-singular solutions of the non-linear coupled eld equations of nite energy such that their energy density has a spacetime dependence of the form: "(x;t) = "(x vt), where v is the velocity of propagation. In relativistic theories, Lorentz (or Galilean) invariance provides all the kink solutions from the purely static ones. The search for nite-energy static solutions in one-dimensional eld theories is tantam ount to the search for nite action trajectories in a natural dynamical system where, the x-coordinate plays the rôle of time; the eld components transmute to positions in the conguration space, and the eld theoretical density energy becomes minus the mechanical potential energy. No wonder the diculties involved in noting kinks in multi-component scalar eld theories: one faces multi-dimensional mechanical systems where integrability is not ensured.

At a very early stage in the (pre-)history of the subject, a (1+1)-dimensional eld theoretical model with two real scalar elds became relevant. Montonen and Sarker-Trullinger-Bishop proposed the deformation of the O(2)-linear sigma model with a potential energy density of U ($_1$; $_2$) = $\frac{1}{2}$ [($_1^2$ + $_2^2$ 1) 2 + $_2^2$], see [2]. It was clear that the zeroes of the potential are two points and hence the hunt for kinks started im m ediately¹. Using a trial orbit m ethod in the associated two-dim ensional m echanical system, Rajaram an identi ed two di erent topological kinks joining the two vacua of the system that live on a straight line and half-ellipses respectively. Only one component of the scalar eld is non-zero in the rst case, but the two-components dier from zero in the second kind of solution; for this reason, these solitary waves are referred to as TK1 (straight line) and TK2/TK2* (upper/lower half-ellipse) kinks in the literature that appeared later. Rajaram an also found one kink associated to a closed trajectory starting from and ending at the same point of the the vacuum orbit. Magyari and Thomas [3] realized that the mechanical system associated with the MSTB model is integrable -there is a second invariant in involution with the mechanical energyand used this fact to show that there exists a whole family of two-component nontopological kinks (NTK2), all of them degenerated in energy with Rajaram an's NTK2 kink; explicit kink form factors were only described by num erical methods.

The main breakthrough in analytically noting all the solitary waves of the M STB model emerged in Reference [4]. Ito discovered that the mechanical problem was not only integrable but that it was H am ilton-Jacobi separable by using elliptic coordinates. In this setting, he showed the analytic formulas for the kink orbits and the kink form factors, unveiling the mathematical reasons for the previously observed striking kink sum rule. Im mediately, the stability of this degenerate kink family was questioned; application of the Morse index theorem solved this problem in [5]. A parallel with the Morse theory of geodesics was established somewhat later in Reference [6]. Thus, a clear connection arose between solitary waves, their stability, and dynamical systems. In Reference [7], several of us showed that the MSTB model is not unique in this respect; two (1+1)-dimensional eld theoretical models with two real scalar elds - referred to as model A and B in that paper-have manifolds of solitary waves with

 $^{^1}$ W e shall refer to the zeroes of the potential as vacua throughout the paper, anticipating their rôle in the quantization of this classical eld theory. A lso, because these two points are related by the internal sym m etry group Z_2 Z_2 generated by $_1$! $_1$ and $_2$! $_2$, we shall som etim es refer to this set as the vacuum orbit.

sim ilar structures. To nd the analytic expression for the kinks of model A, we were prompted to solve an integrable dynamical system classied as Liouville Type I, see [8]. The system belongs to the same class as that found in the MSTB model—the two-dimensional Gamier system [9]—but there are three dierences: (a) the potential energy density is a polynomial of sixth order in the elds (instead of fourth); (b) the vacuum orbit has vepoints (instead of 2), and (c) there are many more stable kinks than in the MSTB model. Model B is characterized by a fourth-order potential energy density in the two scalar elds. The main feature, however, is the need to solve a Liouville integrable system of Type III, i.e. Ham ilton-Jacobi separable in parabolic coordinates. The vacuum orbit has four points and there are manifolds of stable and unstable kinks.

In recent years, all this work has proved to be fruitful in the fram ework of supersym m etric theories. In the dim ensional reduction of a generalized W ess-Zum ino m odel with two chiral super elds, Bazeia-Nascim ento-Ribeiro-Toledo (henceforth referred to as the BNRT model) [10] found one one-component topological kink (TK1) and one two-com ponent topological kink (TK2). In this case, the vacuum orbit has four points and the potential energy density is a polynom ial in the elds of order four. Understanding the BNRT model as a deform ation of model B, some of us discovered the whole manifold of kink orbits [11]. There is kink degeneracy, also found slightly earlier by Shiffm an and Voloshin in one of the topological sectors [12], and, for two critical values of the coupling constant, analytic form ulas for the kink form factors are available. One of them corresponds exactly to model B; the other one leads to a Liouville system of Type IV, Hamilton-Jacobi separable in Cartesian coordinates. Interesting consequences have been translated to the dynamics of intersecting branes [13]. How thick walls grow from one-component kinks is well known. Composite kinks give rise to a non-trivial low energy dynamics for intersecting walls as geodesic motion in the kink m oduli space (the space of the integration constants with a metric inherited from the eld theoretic kinetic energy). A nother supersym m etric m odel that shows a rich pattern of kink solutions is the Wess-Zum ino model itself. The BPS kink states of this N = 2 supersymmetric (1 + 1)D model with a complex scalar eld and holomorphic superpotential were discovered by Vafa et al. in [14]. In [15], two of us studied this system from the point of view of the real-analytic structure. The vacuum orbit having been identied, the ow between the vacuum points was determined as the gradient of the real(im aginary) part of the superpotential. Thus, kink orbits are idential ed with real algebraic curves.

Here, we continue to struggle with the extension of these studies to eld theoretical models with three real scalar elds. In [16], some of us explored the generalization of the MSTB model. The solution of the three-dimensional Garnier system using three-dimensional Jacobi coordinates revealed the existence of an extremely complex variety of kinks. Nevertheless, the structure of the kink manifold and its stability was completely unraveled in [17]. The main goal of the present paper is to identify the kink manifold arising in a family of three-component relativistic eld models with a vacuum manifold that contains several elements or points. This family can be interpreted as the natural generalization of the generalized MSTB model studied in [16][17] in the sense of Stackel-type systems. The most interesting feature of this generalization is that the

number of elements in the vacuum manifold depends on the range of relative values of the coupling constants. Therefore, we can not dierent submodels of our system, which have a very rich structure of kink manifolds. When the energy density of these kinks or solitary waves is studied we not that several families of these solutions are degenerate, which allows us to claim that some kink families indeed consist of more basic kinks, such that their energy density displays several lumps associated with the basic kinks. In our model, we are able to not solutions with two, three or four lumps.

The organization of the paper is as follows: In Section 2 we introduce the model, writing the expressions in Stackel form and describing the dierent spontaneous symmetry-breaking scenarios. Section 3 is divided into four sub-sections. In 3.1, we identify rst-order dierential equations satisted by the kink solutions, reproducing the Bogom ol'nyi procedure in this context. Sub-section 3.2 contains the resolution of these equations. In 3.3, we determ ine the regions where the solutions live and, nally, in Sub-section 3.4, some general comments about the determination of the stability of the kink solutions are overed. In Section 4 we describe the behaviour of solitary wave families in one of the regimes of the model, at the same time discussing their stability properties. Finally, in Section 5 we address some points concerning the dierent extensions of the model.

2 The model

We focus our attention on the search for kink solutions arising in three-com ponent scalar eld models in a (1+1) M inkowskian space-time, whose dynamics is governed by the action functional

$$Z$$
 " #
 $S[] = d^2x \frac{1}{2} \sum_{j=1}^{X^3} Q_{j} Q_{j} U()$;

where we use E instein's convention for G reek indices with the usual metric $_{11} = _{22} = _{1}$, $_{12} = _{21} = _{0}$, and where U () is a smooth non-negative function that depends on the three-component scalar eld $= (_{1};_{2};_{3})$. We use natural units, hence $c = _{1}$, and we shall henceforth denote x^{0} t and x^{1} x. The Euler-Lagrange equations in this case are written as the following system of second-order partial dierential equations

$$\frac{\theta^{2}_{i}}{\theta t^{2}} \quad \frac{\theta^{2}_{i}}{\theta x^{2}} = \frac{\theta U}{\theta_{i}} (_{1};_{2};_{3}) \quad i = 1;2;3:$$
 (1)

K inks are nite-energy solutions of (1), such that the timedependence is dictated by the Lorentz invariance: $_{\rm K}$ (t;x) = $(\frac{{\rm K}^{\rm X}-{\rm V}^{\rm L}}{1-{\rm V}^{\rm 2}})$, and they can be interpreted as extremals of the positive semiled nite energy functional

$$E[] = dx "(x) = dx \frac{1}{2} X^{3} \frac{\theta_{i}}{\theta_{x}} \frac{\theta_{i}}{\theta_{x}} + U(_{1};_{2};_{3}) ; \qquad (2)$$

which maintains this functional nite: $E[\]<+1$, see [1]. Therefore, solitary waves must comply with the asymptotic conditions

a)
$$\lim_{x = 1} 2M$$
 b) $\lim_{x = 1} \frac{d}{dx} = 0$; (3)

where M is the set of zeroes or absolute m in in a of the potential term -that is, M = $f(_1;_2;_3) = R^3 = U(_1;_2;_3) = 0$ g - which are usually referred to as vacua of the theory because the elements of M play this rôle in the corresponding quantum theory.

The usual procedure for tackling the search for kinks in this kind of theory is to interpret (2) as the action functional of a mechanical system in which we think of the variable x as \times in e"; as the coordinates of a unit-m ass point particle, and V=U as the potential function. From this point of view, (1) are merely equations of motion in the new system. In reference [16], the authors deal with the model involving the potential function

$$U(_{1};_{2};_{3}) = \frac{1}{2}(_{1}^{2} + _{2}^{2} + _{3}^{2} 1)^{2} + \frac{1}{2}(_{2}^{2})^{2} + \frac{1}{2}(_{3}^{2})^{2}$$
(4)

and show that the mechanical analogue is not only completely integrable but also H am ilton-Jacobi separable by using a system of three-dimensional elliptic coordinates. In [17], the stability properties of kinks are analyzed and a new approach to search for kinks based on the Bogom ol'nyi decomposition are given in the above system. The authors prove the equivalence between the H am ilton-Jacobi equation and the Bogom ol'nyi approach. The potential function (4) has two zeroes, v = (1;0;0); $v^+ = (1;0;0)$. Therefore, the kinks in this model can be classified into topological and non-topological kinks according to whether the solution connects two different vacua (open orbits) or the solution departs and arrives at a vacuum (closed orbits).

The search for new integrable models is not an easy task. In this sense, we would remark the following quotation from Jacobi in his \Vorlesungen uber D ynam ik", which allows us to see the issue from a dierent perspective: \The main diculty in integrating given dierential equations is to introduce suitable variables which cannot be found by a general rule. Therefore, we must go in the opposite direction and, after nding some remarkable substitution, bok for problems to which it could be successfully applied".

The goal of this paper is to generalize the above model, focusing our attention on models with a greater-than-two number of elements in M , such that we can nd a more sophisticated symmetry-breaking scenario and a richer plethora of solitary waves than before.

Using the same notation as in the reference [16], we now introduce a system of Jacobi elliptic coordinates = ($_1$; $_2$; $_3$), with constants $_3^2 = 1$ $_3^2$, $_2^2 = 1$ $_2^2$ and 1, which is defined as:

$$\begin{array}{rcl}
\frac{2}{1} & = & \frac{1}{2 \cdot 2 \cdot 3} (1 \quad 1)(1 \quad 2)(1 \quad 3) \\
\frac{2}{2} & = & \frac{1}{2(\frac{2}{3} + \frac{2}{2})} (\frac{2}{2} \quad 1)(\frac{2}{2} \quad 2)(\frac{2}{2} \quad 3) \\
\frac{2}{3} & = & \frac{1}{2(\frac{2}{2} + \frac{2}{3})} (\frac{2}{3} \quad 1)(\frac{2}{3} \quad 2)(\frac{2}{3} \quad 3);
\end{array} \tag{5}$$

in which the range of the coordinates is:

$$1 < {}_{1} < {}_{2} < {}_{2} < {}_{2} < {}_{3} < 1 :$$
 (6)

It should be noted that this coordinate transform ation is invariant under the group $G = Z_2^3$ generated by $a! (1)^{ab} a; b=1;2;3.$

Invoking (5), the energy functional can be written as

$$E[] = \frac{Z}{dx} \left(\frac{1}{2} \frac{X^{3}}{g_{jj}} \left(\frac{\partial j}{\partial x} \frac{\partial j}{\partial x} + U(_{1};_{2};_{3}) \right) \right)$$
(7)

where the metric coe cients g $_{jj}$ () = $\frac{1}{4} \frac{f_{j}()}{(j-1)(j-\frac{2}{3})(j-\frac{2}{3})}$ have been introduced.

Here, we set
$$f_j() = Y^3$$

$$(j k = 1 \text{ if } j$$

In the new variables, the potential (4) is written as:

$$U() = \frac{1}{2} \frac{X^3}{\sum_{i=1}^{2} \frac{2_i(i - \frac{-2}{2})(i - \frac{-2}{3})}{f_i(i)}};$$
 (8)

and their zeroes v y v^+ are mapped to one point v $\begin{pmatrix} v \\ 1 \end{pmatrix}$; $\begin{pmatrix} v \\ 2 \end{pmatrix}$; $\begin{pmatrix} v \\ 3 \end{pmatrix}$ $\begin{pmatrix} v \\ 2 \end{pmatrix}$; $\begin{pmatrix} v \\ 3 \end{pmatrix}$ in the elliptic space because of the above-mentioned invariance.

In order to generalize expression (8), we introduce the following potential function

$$U(;^{2}) = X^{3} U_{i}(;^{2}) = \frac{1}{2} X^{3} \frac{2(i - \frac{2}{2})(i - \frac{2}{3})(i - \frac{2}{3})}{f_{i}(;^{2})};$$
 (9)

which becomes a polynomial function of eighth degree in the original elds. Notice that we have added a new factor ($_{i}$ $_{i}$ $_{i}$ $_{i}$ to each of the sum mands in (8). Thus, (9) introduces new degenerate vacua in M, which for $xed \frac{-2}{2}$ and $\frac{-2}{3}$ depend upon the value of the coupling constant $_{i}$ = 1 $_{i}$. Therefore, new scenarios of spontaneous sym metry-breaking and a richer kink manifold arise in this model. Taking into account the range (6) for the elliptic coordinates and form ula (9), we can observe that the new structure of the set M depends on the relative values between the constant $_{i}$ and the xed constants $_{i}$ and 1. For instance, for $_{i}$ > 1 the new factor ($_{i}$ $_{i}$ $_{i}$ does not vanish for any value of $_{i}$ and therefore M has the same structure as that in model (8). However, for $_{i}$ 2 ($_{i}$ 2; $_{i}$ 2) we not new vacua located at the points ($_{i}$ 2; $_{i}$ 2; $_{i}$ 2) and (0; $_{i}$ 2; $_{i}$ 2).

We shall now introduce dierent scenarios for our model depending on the value of the constant 2 . We shall distinguish the number of vacua in each case.

Regime E1: As mentioned above, for 2 2 (1;1) there exists only one vacuum in the elliptic space, minimizing the potential function: v_1 = (0; $_3^2$; $_2^2$). We have a similar situation if the constant 2 takes the discrete values 0, $_3^2$; $_2^2$ or $_2^2$. For this reason we dene the set $L_0 = f0$; $_3^2$; $_2^2$ g [(1;1), taking into account that if $_2^2$ 2 L_0 our model only has a vacuum, $_3^2$, in the elliptic space. In the Cartesian space, the vacuum manifold M can be regarded as the orbit generated by the action of the group $G=H_1$ over the vacuum V_1 , where $V_1=1$ are therefore two vacua in the Cartesian space $V_1=1$ invariant. There are therefore two vacua in the Cartesian space $V_1=1$ and $V_2=1$ invariant.

The kink solutions in this model display the same behaviour as those of the model studied in [16], although the explicit expression of the equations of motion is more complicated because we have a polynomial of degree eight in the original elds. Owing to this similarity, we shall not deal with this regime in our study.

Regime E 2: We now consider the range 2 2 $L_1 = (0; \overline{}_3^2)$ for the coupling constant. In this regime, new zeroes of the potential arise on the plane $_1 = ^2$, in the elliptic space that corresponds to the ellipsoid $\frac{^2}{1} \frac{^2}{2} + \frac{^2}{\frac{^2}{2} \frac{^2}{2}} + \frac{^2}{\frac{^2}{3} \frac{^2}{3}} = 1$ in the Cartesian space. In fact, two vacua arise in the elliptic space, $^{v^1} = (0; _3^2; _2^2)$ and $^{v^2} = (^2; _3^2; _2^2)$, both invariant under the subgroup $H_2 = H_1$. Correspondingly, there are four vacua in the Cartesian space that correspond to the orbit $^2_{i=1}(G=H_i)v_i$. Therefore, we have $M_1 = f^{v_1} = (1;0;0); ^{v_2} = (3;0;0)g$, as depicted in Figure 1.

It is interesting to remark that the range of values 2 2 (1;0) is formally analogous to that in which 2 2 L_1 , interchanging the rôles of the ellipsoids $_1$ = 0 and $_1$ = 2 in the previous reasoning. We shall therefore focus our attention on the range of values L_1 .

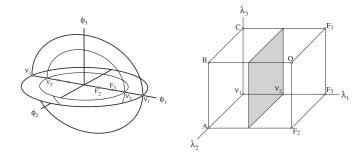


Figure 1: Vacuum manifold in the Cartesian and elliptic spaces: Regime E 2. F_1 , F_2 and F_3 stand for the foci of the ellipsoid; B and C are the extremes of the minor sem i-axis, and A represents the umbilical points.

Regime H 1: In this case, 2 2 $L_2 = (^2_3; ^2_2)$. From the values of 2 and the range of the elliptic coordinates, the zeroes of the potential term (9) arise on the plane $_2 = ^2$, which is equivalent to the hyperboloid of one sheet $\frac{^2}{1^{-\frac{1}{2}}} + \frac{^2}{2^{-\frac{2}{2}}} = 1 + \frac{^2}{2^{-\frac{2}{3}}}$ in the Cartesian space. We not three vacua located at the points $^{v^1} = (0; ^2_3; ^2_2)$, $^{v^2} = (^2_3; ^2_3; ^2_2)$ and $^{v^3} = (0; ^2_3; ^2_2)$. The vacua v_1 and v_2 remain invariant under v_1 , whereas v_2 is invariant under v_3 and v_4 are eight vacua in the Cartesian space corresponding to the orbit v_3 is invariant under v_4 and v_5 in the coordinates v_4 in v_5 in section 4, for the sake of clarity we shall focus on this regime in order to describe in detail a particular kink manifold of the model instead of discussing it in each single regime.

Regime H 2: This case is characterized by 2 2 L₃ = $\binom{2}{2}$;1). Applying the same reasoning as before, we not that new vacua arise on the plane $_3$ = 2 ;

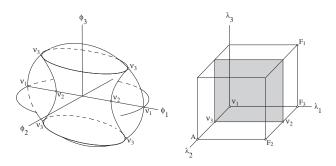


Figure 2: Vacuum manifold in the Cartesian and elliptic spaces: Regime H 1.

that is, the hyperboloid of two sheets $\frac{\frac{2}{1}}{1} = 1 + \frac{\frac{2}{2}}{2} + \frac{\frac{2}{3}}{2} = \frac{2}{3}$ in Cartesian coordinates. In particular, the potential has four minima: $v_1 = (0; \frac{2}{3}; \frac{2}{2}); v_2 = (\frac{2}{3}; \frac{2}{2}; \frac{2}{2}); v_3 = (0; \frac{2}{2}; \frac{2}{2}), \text{ and } v_4 = (0; \frac{2}{3}; \frac{2}{3}).$ The vacuum v_4 is invariant under $v_4 = 1$ 1 $v_4 = 1$ 1

M₃ = f^{v₁} = (1;0;0); ^{v₂} = (;0;0);

$$v_3$$
 = $-$;0; $-$ 3/3 ; v_4 = $-$; $-$ 2/2 ; 0 :

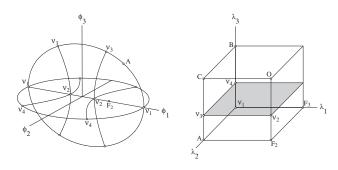


Figure 3: Vacuum manifold in the Cartesian and elliptic spaces: Regime H 2.

Regime H 2^0 : In this case, the coupling constant 2 is set equal to unity. In the internal elliptic space we can read the minima as V_1 = $(0; \frac{2}{3}; \frac{2}{2})$, V_2 = $(\frac{2}{3}; \frac{2}{2}; 1)$, V_3 = $(0; \frac{2}{2}; 1)$ and V_4 = $(0; \frac{2}{3}; 1)$, which correspond to eight minima in the Cartesian space: M $_3^0$ = $\lim_{\substack{2 \ 1 \ 2 \ 1}}$ M $_3$ = f V_1 = (1;0;0); V_2 = (0;0;0); V_3 = (0;0;0;0); V_4 = (0;0;0;0)9.

In this latter case, the plane $_3=1$ is introduced into the elliptic space. Unlike the previously introduced planes, this is no longer a regular one, and this can be readily seen in the degeneracy exhibited by the H 2 vacuum manifold at the lim it 2 ! 1; this singular plane corresponds to the plane $_1=0$ in Cartesian coordinates. Regarding the kink manifold, this is basically the same as that of the H 2 model, except that the kink solutions existing on the two sheets of the hyperboloid and in between them now degenerate into kink solutions on the plane $_1=0$. This situation is the 3D analogue of model A in [7].

3 First-order equations and Kink Manifolds

3.1 The superpotential and the Bogom ol'nyi arrangem ent

We notice that the potential (9) determ ines a Stackel system [8]. Therefore, the Ham ilton-Jacobi equation of the mechanical analogue is separable using the system of Jacobi elliptic coordinates. However, here we shall make use of the Bogom ol'nyi arrangement in order to obtain the kink manifold of our model. The two procedures are equivalent (see [17]) but the second one allows us to identify the supersymmetric extension of our eld theory, given that if the energy functional (7) can be written as

$$E[] = \frac{Z}{dx} \frac{1}{2} \frac{X^{3}}{g_{jj}} \frac{Q_{j}}{Qx} \frac{1}{g_{jj}} \frac{QW()}{Q_{j}}^{2} \frac{Z}{dx} \frac{dW()}{dx}$$
(10)

for some function W ($_{i}$), then the underlying eld theory has a supersym m etric extension in which the function W plays the rôle of superpotential in the supersym m etric eld theory, see [18]. Therefore, the superpotential W must comply with

2U () =
$$g_{11}^{1}$$
 () $\frac{@W}{@_{1}}^{2}$ + g_{22}^{1} () $\frac{@W}{@_{2}}^{2}$ + g_{33}^{1} () $\frac{@W}{@_{3}}^{2}$:

Plugging the expression of the potential function (9) and the metric coecients into the above equation, we have

$$\frac{X^{3}}{\underbrace{\frac{2}{i}(\underline{i} \quad 2)^{2} \underbrace{\frac{Q}{j=2}(\underline{i} \quad -2)}_{j=2}}_{\underline{i=1}} = \frac{X^{3}}{\underbrace{\frac{4(\underline{i} \quad 1)^{Q} \underbrace{\frac{3}{j=2}(\underline{i} \quad -2)}_{j=2}}_{\underline{f_{i}}(\underline{i})}} \underbrace{\frac{0W}{\underline{0}_{\underline{i}}}^{2}}_{\underline{i}}^{2};$$

which can be solved easily by the ansatz $W = W_1(_1) + W_2(_2) + W_3(_3)$. The three resulting decoupled ordinary di erential equations

$$\frac{dW_{i}}{d_{i}}^{2} = \frac{{}_{i}^{2}(_{i}^{2})^{2}}{4(1_{i}^{2})} ; i = 1;2;3$$

lead us to the expression of the superpotential function W

$$W^{(1;2;3)}() = X^{3} \qquad W_{i}() = \frac{1}{15} X^{3} \qquad (1)^{i}P_{2}()^{i}P_{2}()^{i} \qquad ; \qquad i = 0;1;$$

where $P_2(_{i}) = 2d + d_{i} = 3^{2}$, with $d = (5^{2} = 4)$.

Extrem altra jectories for the energy functional (10) arise if the following system of rst-order di erential equations

$$\frac{d_{i}}{dx} = (1)^{i}g_{ii}^{1}()\frac{dW_{i}}{d_{i}}
= (1)^{i}2\frac{i(i^{2})(i^{2})(i^{2})(i^{2})}{f_{i}()} \frac{-2}{3}p_{i}$$
(11)

where $_{i}$ = 0;1 and $_{i}$ = 1;2;3 is satis ed, because the squared term s in (10) are always positive and the last one is a constant. Due to the indeterm inacy of the signs $_{1}$, $_{2}$ and $_{3}$, (11) constitutes eight system s of ordinary di erential equations. Nevertheless, this set of system s is easier to solve than second-order (Euler-Lagrange) equations. In order to obtain a complete kink solution we have to join solutions from the rst-order di erential equations with dierent choices of the signs ($_{1}$) in dierent intervals covering the real line. The reason for this is that the rst-order dierential equations inherit the information of the second-order equations dended piecewise. Assuming that we search for continuous and dierentiable solutions, the sequence of signs ($_{1}$) is corresponding to the dierent pieces that constitutes a solution is prescribed. In section 4 we shall illustrate this approach in several cases. From (10) it is readily seen that the energy of a solitary wave, solution of (11) with only one piece, depends only on the topological charge of the solution. In this case, it is said that the Bogom ol'nyi bound is saturated. However, if the orbit is given by = $_{1}^{3}$, where J is the number of pieces of and is stands for the jth piece, we have:

$$E[] = X Z dx \frac{dW()}{dx} = X Z X^{3} \frac{\partial W^{f ig_{j}}}{\partial i} d_{i}$$

$$= X^{J} (W^{f ig_{j}}(^{j}_{nal}) W^{f ig_{j}}(^{j}_{initial})); \qquad (12)$$

where $f_{i}g_{j}$ represents the values of the i param eters for the jth piece of the solution.

3.2 Solutions via quadratures

In order to solve system (11), we rewrite it in the form:

$$\frac{d_{i}}{(1)^{i}2^{p}} \frac{d_{i}}{1_{i}} \frac{d_{i}}{d_{i-1}} (i_{i} c_{i}) = \frac{dx}{f_{i}(1)} ; i = 1;2;3;$$
 (13)

where we have de ned c = (2 ; 2_2 ; 2_3 ; 2_4) and 2_4 = 0 for future convenience. The sum of these equations gives

$$X^{3} = \frac{d_{i}}{(1)^{i}2^{p} \frac{d_{i}}{1} \frac{d_{i}}{d_{i}} \frac{d_{i}}{d_{i}} = 0} = 0$$
 (14)

Multiplying each side of (13) by $_{i}$ and sum ming over i, we obtain:

$$X^{3} = \frac{id}{(1)^{i}2^{p} \frac{id}{1}} \underbrace{\frac{i}{i}}_{i} \underbrace{\frac{i}{i}}_{j=1} \underbrace{(i \quad c_{j})}_{j=1} = 0 :$$
 (15)

A lso, multiplying (13) by $^2_{\rm i}$ and sum m ing again over iwe reach the equation that establishes the dependence of the kink components on x

$$\frac{X^{3}}{\sum_{i=1}^{3} \frac{p^{2} d^{i}}{(1)^{i} 2^{p} \frac{1}{1} \sum_{i=1}^{3} \frac{q^{i}}{q^{i}} (1)^{i}} = dx :$$
(16)

We shall now determ ine the kink orbits and the form factor by invoking (14), (15), and (16). Integration of the 1st two equations,

$$\frac{X^{3}}{\sum_{i=1}^{2} \frac{(1)^{i}}{2}} \frac{Z}{P} \frac{\frac{d}{1}}{\frac{1}{\sum_{i=1}^{4} (i)} \frac{C_{i}}{\sum_{j=1}^{4} (i)}} = 2$$

$$\frac{X^{3}}{\sum_{i=1}^{2} \frac{(1)^{i}}{2}} \frac{Z}{P} \frac{\frac{id}{1}}{\frac{1}{\sum_{i=1}^{4} (i)} \frac{C_{i}}{\sum_{j=1}^{4} (i)}} = 3$$

leads us to the expression of the generic kink orbits:

$$e^{2} = Y^{4} \quad p \xrightarrow{\frac{1}{1}} p \xrightarrow{\frac{1}{1}} p \xrightarrow{\frac{1}{1}} c_{j} \qquad Y^{4} \quad p \xrightarrow{\frac{1}{2}} p \xrightarrow{\frac{1}{1}} c_{j} \qquad Y^{4} \quad p \xrightarrow{\frac{1}{1}} p \xrightarrow{\frac{1$$

where $F_j(c) = \begin{pmatrix} p & Q_4 \\ 1 & c_j \end{pmatrix}_{l=1; l \in j} (c_j & c_l)$, and $_2$ and $_3$ are arbitrary real constants that specify a particular kink orbit.

The integration of (16)

$$\frac{X^{3}}{\sum_{i=1}^{2} \frac{(1)^{i}}{2}} \frac{Z}{P} \frac{P}{1} \frac{Q_{i}^{2} Q_{i}^{2}}{Q_{j=1}^{2} (i C_{j})} = 1 + x$$

gives us the form factor of the kink:

$$e^{2(1+x)} = y^{3} p_{\frac{1}{1}} p_{\frac{1}{1$$

 $_{1}$ being an integration constant associated with the translational invariance of the system . Expressions (17), (18) and (19) provide us with the whole manifold of solitary waves.

3.3 Frontiers and barriers. Basic kinks

We shall now prove that the set of solitary waves is connect to living in a bounded region of the internal space, which in fact corresponds to a parallelepiped in the elliptic space. For the sake of clarity, we shall restrict our study to the range 2 2 L, where L = $^3_{i=1}$ L_i is the set in which the kink manifold is richest, see Section 2. This include the regimes E 2, H 1, and H 2. Squaring the rst equation in (13), and dening the generalized momentum $_1 = g_{11}$ () $\frac{d_1}{dx}$, we have:

$$\frac{1}{2} \, {}_{1}^{2} \, \frac{{}_{1}^{2} \left(\, {}_{1} \, {}_{2} \right)^{2}}{8 (1 \, {}_{1})} = 0; \tag{20}$$

Equation (20) can be regarded as that governing the motion of a particle moving under the in uence of the potential function

$$U(_{1}) = \begin{cases} 8 & \frac{_{1}^{2}(_{1} & _{2})^{2}}{8(1 & _{1})} \\ \vdots & \vdots & \vdots \\ 1 & \vdots & \vdots \\ \end{cases}; \quad 1 < _{1} < _{2}^{2} < _{1} < 1$$

The function has at least one minimum in $_1=0$ and a second one in $_1=^2$ if $_2^2$ L $_1$. Furthermore, the function U ($_1$) goes to $_1$ as $_1$ tends to $_1$. Thus the bounded motion can only occur in the interval $[0;\frac{2}{3}]$. This, combined with the boundary conditions, leads us to the conclusion that the kink solutions lie in the parallelepiped P₃(0) = $[0;\frac{2}{3}]$ $[\frac{2}{3};\frac{2}{2}]$ $[\frac{2}{2};1]$.

There is still more inform ation that can be extracted following this procedure, owing to the appearance of a second minimum. Let us rst x a value 2 in L, and let us set 2 2 L $_i$ for some i that depends on 2 . Squaring the i^{th} equation of the system (13) and dening the generalized momentum $_i = g_{ii}()\frac{d_i}{dx}$, we arrive at a similar one-dimensional dynamics:

$$\frac{1}{2} \stackrel{?}{=} \frac{\binom{2}{i} (\frac{i}{i} - \binom{2}{i})^2}{8(1 - i)} = 0:$$
 (21)

A coordingly the corresponding potential function is now de ned by U ($_1$) if i = 1 and $_8$

$$U(_{i}) = \begin{cases} \frac{2}{i}(_{i} - _{i}^{2})^{2} \\ 8(1 - _{i}) \end{cases} ; m infL_{i}g < _{i} < m axfL_{i}g$$

$$\vdots \qquad \vdots \not \geq L_{i}$$

for i=2;3. The minimum $_{i}=^{2}$ now separates the bounded motion of the one-dimensional system into two intervals – the $_{i}$ 2 $L_{i}=$ [m infL $_{i}$ g; 2] interval and the $_{i}$ 2 $L_{i}^{+}=$ [2 ;m axfL $_{i}$ g] interval –, and into the trivial motion $_{i}=L_{i}^{0}=^{2}$. This, together with the asymptotic conditions, leads us to conclude that, besides living in $P_{3}(0)$, the kink solutions lie entirely in the sets

$$P_3(0)^{,0;+} = f 2 P_3(0) \text{ with } i 2 L_i^{,0;+} g$$
:

This decomposition of the parallelepiped $P_3(0)$ is, for the case we shall study in detail, regime H 1, as follows (see Figure 2):

$$P_3(0) = P_3(0) [P_3(0)^0 [P_3(0)^+ = [0; \frac{2}{3}] [\frac{2}{3}; \frac{2}{3}] [\frac{2}{2};1][$$

$$[[0; \frac{2}{3}] f^2g [\frac{2}{2};1][[0; \frac{2}{3}] [\frac{2}{2};\frac{2}{2}] [\frac{2}{2};1]:$$

The parallelepipeds $P_3(0)$ and $P_3(0)^+$ contain families of solutions that depend on two and three parameters, whereas the plane $P_3(0)^0$ only contains two-parametric solutions.

Thus, introduction of the factor $(i^2)^2$ into the potential function U () leads us (within our range of study) to a new con nem ent of kink solutions in the parallelepiped

 $P_3(0)$. The generic kink solutions divide into two sectors and, in addition to this, a new kind of two-parametric solutions arises: those satisfying $_i = ^2$. Consequently, the kink manifold can be decomposed as follows:

$$C = C_i + C_i^0 + C_i^+;$$
 (22)

where $C_i^{0,t}$ represent the class of kink solutions with i 2 , i = 2 and i respectively.

3.4 Stability

In this sub-section we discuss how to determ ine the stability properties of the kink solutions. For the whole variety of kink solutions in this system, it is not possible to solve $_1$; $_2$ and $_3$ in terms of elementary functions of x. Therefore, it is not possible to explicitly write out the Hessian operator for any kink in the model and, hence, the stability properties cannot be studied through analysis of its spectrum.

To determ ine the stability of the solutions, we use instead the arguments developed in Ref. [16] based on the Jacobi elds along kink solutions. A lthough the treatment depicted in that paper is for a deformed Sigma O (3) model, the extension to this model can be readily carried out. Following this procedure, a rule establishing the stability (instability) of the solutions is obtained: each solution crossing either the edge F_1F_3 f_3^2 ; g_3^3 ; g_3 or the edge g_3 g_3 g_3 becomes an unstable solution, since these two edges constitute lines of conjugate points of each vacuum of the theory.

The key point is that the superpotential function is not di erentiable over either of these two edges and, consequently, the energy of the kink (12) is not a topological quantity since it depends on the value of the superpotential at the crossing points.

In what follows, and bearing this remark in mind, we shall only mention the character of each of the kinks described.

4 Description of the Kinkmanifold in the H 1 regime

The description of the kink manifold in the dierent regimes arising in our model is a long and tedious task. We shall therefore focus our attention on a particular example: the H 1 regime. Nevertheless, this case will suce to illustrate the general features that also arise in other regimes of our model. We shall now describe the behaviour of the kinks that arise in the H 1 regime of our model. We can not basic kinks, similar to the solutions TK 1 and TK 2 in MSTB model, that are placed on the edges of the characteristic parallelepiped in the elliptic space (see gures 4,5 and 6). These solutions are the simplest kinks in our model and they consist of a single lump, such that they can be interpreted as an extended particle. We shall show that the kink manifold includes other kink solutions involving several lumps associated with the basic kinks.

We recall some remarkable points of the H 1 regime from the previous sections: The number of minima is three in the \elliptic" space, and eight in the Cartesian one (see Figure 2): M $_2$ = $_{v_1}^{v_2}$ = ($_{1;0;0}$); $_{v_2}^{v_2}$ = ($_{3;0;0}$); $_{3;0}^{v_3}$ = ($_{3;0;0}^{-3;0;0}$).

The ellipsoid E $\frac{2}{1} + \frac{2}{\frac{2}{2}} + \frac{3}{\frac{2}{3}} = 1$ (that is, = (0; 2; 3)), the one-sheet hyperboloid H $\frac{2}{1+2} + \frac{2}{\frac{2}{2}-2} + \frac{3}{2-\frac{2}{3}-2} = 1$ (or $_2 = _2^2$), and the planes $_{2,3} = _3^2$

Finally, we describe the kink manifold in these cases. We can distinguish: A, Non-generic, two-parametric families, and B, Generic, three-parametric families of solitary waves:

A Two-Param etric families of solutions:

A 1 Solutions on the ellipsoid E.

The potential term U_1 vanishes on this surface. A coordingly, the superpotential function is:

$$W^{(2;3)}(_{2;3}) = X^3 W_{i}(_{i}) = \frac{1}{15} X^3 (_{1})^{i} P_{2}(_{i})^{p} \overline{1}_{i} ; i = 0;1:$$

The orbit of these solutions is given by

$$e^{2^{2}} = \underbrace{\frac{y^{3}}{p} \frac{p}{1} \frac{p}{2} + p \frac{p}{1} \frac{c_{j}}{1}}_{j=1} \underbrace{\frac{(\frac{1}{2})^{2}c_{j}}{p}}_{j=1} \underbrace{\frac{p}{1} \frac{p}{1} \frac{p}{3} + p \frac{p}{1} \frac{c_{j}}{1}}_{j=1} \underbrace{\frac{(\frac{1}{3})^{3}c_{j}}{p}}_{j=1} ;$$

- 2 being an arbitrary real constant. We have two kind of solutions:
- i) $T_E^{v_1\,v_3}$: Stable topological solutions that connect the m in im a v^1 and v^3 after having crossed the plane $_1$ = 0.
- ii) N $_{\rm E}^{\rm V_3}$: U nstable non-topological solutions that join them in im um $\rm v^3$ w ith itself. The trajectory of these solutions starts from $\rm v_3$, reaches the plane $\rm v_1=0$, and -after crossing the um bilical point A returns to the same point $\rm v^3$; see Figure 4.

The energy of these solutions can easily be calculated by integrating dW along their respective orbits:

E [T_E^{v₁,v₃}] =
$$dW = dW^{(0,0)} + dW^{(0,1)}$$

= $dW_2^{v_1,v_3}$ = $dW_2^0 + 2 dW_3^0 = \frac{1}{15} P_2(^2) g_3^2 P_2(^2)$
= $\frac{2}{3} \frac{5}{5} g_3^3 g_3^3 2 \frac{5}{5} g_3^3$

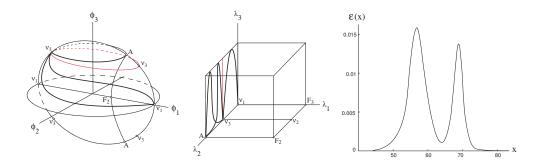


Figure 4: Solitary waves on E in the Cartesian (left) and elliptic (m iddle) spaces. Energy density of a kink of the fam ily $T_{\rm E}^{\rm V_1\,N_3}$ (right).

A 2 Solutions on the plane $_3 = 0$.

In this case, the term $s U_2$ and U_1 of the potential vanish, but not simultaneously. The former vanishes over $_2=_3^2$, and the latter over $_1=_3^2$. Because of this, two superpotential functions appear, and hence two systems of dierential equations must be involved in order to determ ine this solution. Nevertheless, we can synthesize W as follows:

$$W^{(k;3)}(k;3) = \frac{1}{15} X_{i=k;3} (1)^{i} P_{2}(i)^{p} \frac{1}{1} i ; i = 0;1;$$

where k=1 for $_2=\frac{2}{3}$, and k=2 for $_1=\frac{2}{3}$. The equations of the orbit on the plane $_k=\frac{2}{3}$ are:

$$e^{2^{2}} = \underbrace{\frac{Y^{4}}{p} \frac{p}{1} \frac{p}{k} \frac{p}{1} \frac{p}{k} \frac{p}{1} \frac{(1) k (c_{j} c_{3})}{p}}_{k} + \frac{p}{1} \frac{p}{1} \frac{p}{k} \frac{p}{1} \frac{p}{1} \frac{p}{3} + \frac{p}{1} \frac{p}{1} \frac{(1) 3 (c_{j} c_{3})}{p}}{p}}_{j = 1} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1} \frac{p}{1} \frac{p}{1} \frac{p}{1}}_{j \in 3}}_{i} : \underbrace{\frac{(1) 3 (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{i} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1} \frac{p}{1} \frac{p}{1}}_{j \in 3}}_{i} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3}}_{j \in 3} : \underbrace{\frac{(1) k (c_{j} c_{3})}{p} \frac{p}{1}}_{j \in 3}}_{j \in 3$$

Again we have two kinds of solutions:

- i) T $_3^{v_1,v_2}$: Unstable topological solutions linking the vacua v^1 and v^2 . These solutions leave v^1 , intersect the axis $_2$ and the segment F_1F_3 consecutively, and nally arrive at v^2 , as depicted in Figure 5.
- ii) N $_{_3}^{v_2}$: U nstable non-topological solutions connecting v^2 . The solutions go from v^2 , intersect the axis $_2$ = 0, cross the focus F_2 , and return to the initial point v^2 .

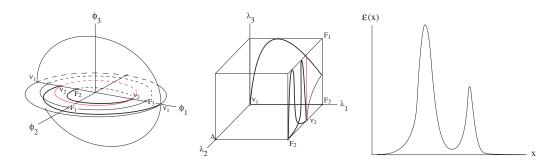


Figure 5: Solitary waves on $_3 = 0$ in the Cartesian (left) and elliptic (middle) spaces. Energy density of a kink of these families (right).

The computation of the energies is as follows:

$$E [T_{3}^{v_{1} v_{2}}] = \frac{2}{3} \frac{5}{5} \frac{3}{5} \frac{1}{5} \frac{2}{5} \frac{2}{5} \frac{3}{2}$$

$$E [N_{3}^{v_{2}}] = \frac{4}{3} \frac{3}{5} + \frac{3}{2} \frac{5}{5} :$$

A 3 Solutions on the plane $_2 = 0$, see Figure 6.

Now, the term $s\,U_3$ and U_2 of the potential vanish over $_3=_2^2$ and $_2=_2^2$, respectively. The two superpotential functions that appear can be synthesized in a similar way:

$$W^{(1;k)}(x_1; k) = \frac{1}{15} X_{i=1;k} (x_1)^{i} P_2(x_1)^{p} \frac{1}{1 - i} ; i = 0;1;$$

where k=2 for $_3=\frac{2}{2}$ and k=3 for $_2=\frac{2}{2}$. The equations of the orbit on the plane $_k=\frac{2}{2}$ are:

We now have three classes of solutions:

- i) $T_{2}^{v_{1},v_{2}}$: Stable topological solutions that join the m in im a v^{1} and v^{2} , as can be observed in Figure 6.
- ii) $T_2^{v_3}$: Unstable topological solutions that connect the point v^3 with the minimum, which is its rejection by the transformation $_3$! $_3$, previously crossing the focus F_3 .
- iii) T $_{2}^{v_{2}N_{3}}$: Unstable topological solutions that link the points v^{2} and v^{3} . In this case, the solutions depart from v^{2} , and nally arrive at v^{3} after intersecting the axis $_{3}$.

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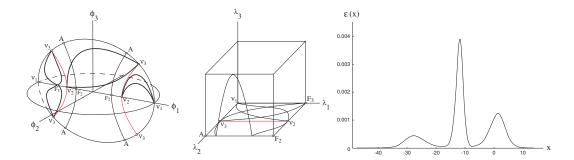


Figure 6: Solitary waves on $_2$ = 0 in the Cartesian (left) and elliptic (middle) spaces. Energy density of a concrete kink of the family $T_2^{v_3}$ (right).

The energies for these solutions are:

$$E \left[T_{2}^{v_{1} N_{2}}\right] = \frac{2}{3} \cdot \frac{1}{5} \quad ^{2} \quad \frac{5}{5} \quad ^{3}$$

$$E \left[T_{2}^{v_{3}}\right] = \frac{4}{3} \cdot \frac{1}{5} \quad ^{2} \quad \frac{5}{5} \quad ^{3} \quad ;$$

providing a simple kink energy sum rule: $2E [T_{2}^{v_1 N_2}] = E [T_{2}^{v_3}]$. The remaining energy is:

$$E \left[T_{2}^{v_{2} N_{3}}\right] = \frac{2}{3} \quad \frac{5}{5} \quad 3 \quad 2 \quad \frac{5}{5} \quad 3 \quad \frac{1}{5} \quad 2$$
:

In gure 6(right), we have depicted the energy density "(x) of a m em ber of the fam ily T $_2^{\rm V_3}$. W e notice that the kinks of this fam ily consist of three basic lum ps.

A 4 Solutions on the hyperboloid.

The term U_2 vanishes over $_2 = ^2$ and hence the superpotential function is:

$$W^{(1;3)}(_{1};_{3}) = \frac{1}{15} X_{_{i=1;3}} (_{1})^{_{i}} P_{2}(_{i})^{_{p}} \overline{1}_{_{i}} ; \quad _{i} = 0;1:$$

The equation of the orbit is:

In this case, only one family is found.

 $T_H^{v_2 N_3}$: The trajectories of these stable solutions connect the points v^2 and v^3 , previously intersecting the plane $_1 = 0$, as is shown in Figure 7. Notice that the energy density in this case comprises two basic lumps.

The energy is:

$$\mathbb{E}\left[T_{H}^{\mathbf{v}_{2}N_{3}}\right] = \frac{2}{3} \quad \frac{\frac{5}{3}}{5} \quad \frac{3}{3} \quad \frac{1}{5} \quad 2 \quad 2 \quad \frac{\frac{5}{2}}{5} \quad \frac{3}{2} \quad :$$

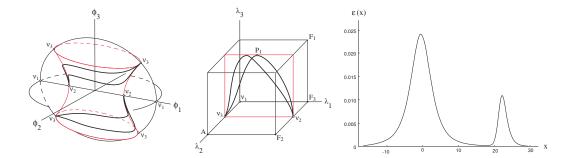


Figure 7: Solitary waves on H in the Cartesian (left) and elliptic (middle) spaces. Energy density of a kink of this family (right).

B Three-Param etric families of solutions. We not three kinds of solutions:

B 1 Solutions located inside the ellipsoid and outside the hyperboloid, see Figure 8:

- i) T v_1,v_2 : Stable topological solutions that join v^1 and v^2 . The solutions em erge from v^1 , later cross the plane $_1$ = 0, and nally arrive at v^2 .
- ii) T^{v_3} : Unstable topological solutions, which start from a minimum v^3 , consecutively cross the planes $_1=0$ and $_3=0$, intersecting the F_1F_3 edge, and nally arrive at v^3 . Notice that the energy density in this case comprises four basic lumps.

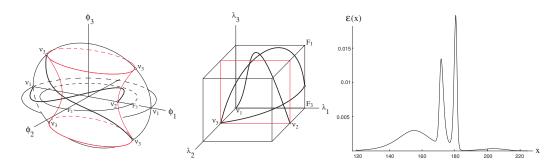


Figure 8: G eneric solitary waves in the Cartesian (left) and elliptic (m iddle) spaces. Energy density of a kink of the fam ily T^{v_3} (right).

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Their energies are:

$$E[T^{v_1 v_2}] = \frac{2}{3} \cdot \frac{5}{5} \quad ^3 \quad \frac{1}{5} \quad ^2 \quad 2 \cdot \frac{\frac{5}{2}}{5} \quad ^{\frac{3}{2}}$$

$$E[T^{v_3}] = \frac{4}{3} \cdot \frac{5}{5} \quad ^3 \quad \frac{1}{5} \quad ^2 \quad \frac{\frac{5}{2}}{5} \quad ^{\frac{3}{2}}$$

B 2 Solutions located inside the hyperboloid:

i) T^{v_2,v_3} : These are unstable solutions. They leave v^2 , cross the plane $_1=0$, later intersect the hyperbola AF $_2$, cross the plane $_1=0$ again, and nally arrive at the point v^3 ; see Figure 9.

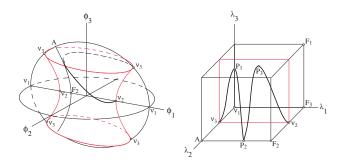


Figure 9: Generic solitary waves in the Cartesian (left) and elliptic (right) spaces.

The energy in this case is:

$$E[T^{v_2 v_3}] = \frac{2}{3} \quad \frac{5}{5} \quad \frac{3}{3} \quad \frac{1}{5} \quad ^2 \quad 2 \quad \frac{5}{5} \quad \frac{3}{2} \quad 2 \quad \frac{5}{5} \quad ^3$$

To complete the previous energy calculations, the kink energy sum rules satis ed by the generic solutions are o ered:

{
$$E [T^{v_1 N_2}] = E [T^{v_1 N_2}]$$
}
{ $2E [T^{v_2 N_3}] = E [N^{v_3}] + E [T^{v_2 N_3}] + E [T^{v_2 N_3}]$ }
{ $2E [T^{v_3}] = E [T^{v_1 N_3}] + E [T^{v_2 N_3}]$ }
3E $[T^{v_1 N_2}]$.

See Sub-section 3.1 of Reference [16] for an explanation of the origin of these rules in a simpler setting. We stress that the decomposition of the kink energy density in several lumps is due to the kink energy sum rules.

Finally, as an example we depict the kink form factor (Fig. 10 and Fig. 11) for the two unstable generic solutions.

5 Further Comments

It is possible to generalize this kind of model in two senses; we enlarge the internal space with N scalar elds and we include a greater number of coupling constants $^2_{i}$.

1. To study the generalization of this kind of system to N dimensions, it is rst necessary to introduce N-dimensional Jacobi elliptic coordinates. An appropriate explanation of these can be seen in [16]. The potential function we propose for the system is as follows:

$$U(;^{2}) = X^{N} \qquad U_{i}(;^{2}) = \frac{1}{2} X^{N} \qquad \frac{{}_{i}^{2}(_{i} \qquad {}_{j=2}^{2}(_{i} \qquad {}_{j}^{2})}{_{j=2}^{2}(_{i} \qquad {}_{j}^{2})};$$

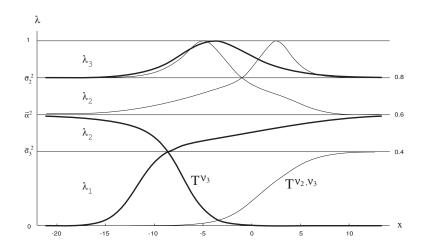


Figure 10: Factor form for the T v_2 v_3 and T v_3 solutions. For the T v_3 solution, we have taken $_1$ = 0, $_2$ = 5 and $_3$ = 5, whereas for the T v_2 v_3 solution the constants are $_1$ = $_2$ = $_3$ = 0.

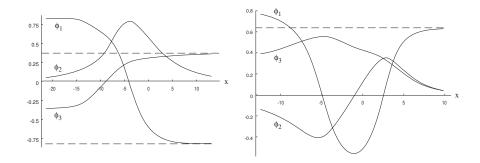


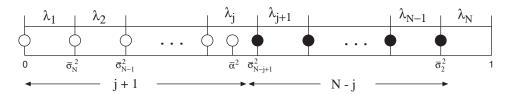
Figure 11: Form factors in the Cartesian space for the T $^{\rm V_3}$ and T $^{\rm V_2}$ $^{\rm V_3}$ solutions

where the coupling constants together with the coordinates satis es the chain

$$1 < {}_{1} < {}_{2} < {}_{2} < ::: < {}_{N} _{1} < {}_{2} < {}_{N} < 1 = {}_{1}^{2}$$
:

The denom inator is $f_i() = \frac{Q_{j \in i}()}{j \in i}()$, and i = 1 is a real positive constant. The function U(i; i) is positive sem i-de nite and presents a number of zeroes, depending on i = 1. The most interesting kink manifold appears when i = 1 is i = 1; i =

We shall now brie y study the vacuum manifold in all the N dierent cases at once. Let us set 2 such that 2 2 L_j for some j between 1 and N. To nd a zero of the function U (; 2), we must make every term U_i (; 2) vanish. To visualize the process, we shall seek help from the following graphic



Each circle in the $_k$ block represents a value that $_k$ can take to make the term $U_k(\ ;\ ^2)$ null. Each value appearing in the vacuum coordinates will be represented by a full circle, and hence each vacuum in the elliptic space is represented by N full circles. To ll the N $_j$ circles to the right of $_j$, there is only one possibility, as seen in the gure, but to ll the remaining $_j$ circles we have a number of dierent ways equal to the number of permutations of $_j+1$ elements, $_j$ of them being repeated. Therefore, we have $P_{j;l}^{(j+1)}=_j+1$ zeroes of the U($_j$) function, $_j$ of them being on the plane $_j=_j^2$. To gure out the number of corresponding Cartesian vacua, we only need to take into account the multiplicity of each elliptic vacuum. By doing this, we conclude that by introducing a regular plane $_j=_j^2$ there are $V=_j^2$ there are $V=_j^2$ cartesian vacua. The kink manifold thus decomposes into V_j^2 disconnected sectors [7].

2. The second generalization considers not only one parameter, 2 , but several of them . The generalized potential is constructed as follows.

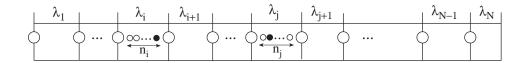
Let us consider num bers $n_i=0$;1;2;:::with i=1;:::;N , and let us de ne $(n_1+\dots+n_N)$ di erent param eters $^2_{ij}$, such that for each $n_i \in 0$, $^2_{ij} \in L_i$ and j=1;:::; n_i . We can therefore construct the N -dim ensional potential:

$$U = \sum_{i=1}^{X^{N}} U_{i}(; i_{ij}) = \frac{1}{2} \sum_{i=1}^{X^{N}} \frac{\frac{2}{i}(i_{2})(i_{2})}{f_{i}(i_{3})} \frac{2}{3} \frac{Y^{i}}{i_{1i}^{i_{1i}}} (i_{2})^{2}; \qquad (23)$$

The case in which $P_{N}^{i=1} n_{i} = 0$ corresponds to the deformed 0 (N) linear sigma m odel [16] and the case $P_{N}^{i=1} n_{i} = 1$, with N = 3, is precisely the m odel studied in the previous sections.

As $\sum_{i=1}^{N} n_i$ increases, the vacuum manifold becomes more and more abundant owing to the appearance of an increasing number of roots in the potential. An easy way to

account for the vacuum manifold V is through the corresponding generalization of the previous graphic



In this picture $(n_1 + :::+ n_N)$ additional circles appear, —holes for short, since every $^2_{ij}$ is easily seen to be a root of the U_i term in (23). Com putation of the number of vacua now proves to be an easy task given that, as before, each vacuum point is represented by N lled circles. It happens that the number of vacua—including v_1 , which corresponds to all the —holes emptied—is given by:

$$C \operatorname{ard}(V) = 1 + \bigvee_{q=1}^{X^{N}} N_{q};$$

where N $_{\rm q}$ is the number of vacuum points with q led -holes, which can be calculated readily using combinatorial techniques.

Regarding the kink manifold, and looking at the corresponding rst-order equations, for each $^2_{ij}$ we can deduce a connement of the solutions in $P_3(0)$ similar to that obtained in section 3. Therefore, a number of $2^{(n_1+\dots+n_N)}$ subsets of $P_3(0)$ that host general kink solutions appear.

The purpose of this construction is now clear. Recalling the stability criterion and the connem ent of the solutions due to the factors $(i_1, i_2)^2$, we can isolate the edges $F_1F_3 = f_3^2$; i_3^3 ; i_3^3 g and i_3^3 g and i_4^3 g. Proceeding in this way, we cannot subsets of the congulation space in which only stable solutions emerge.

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