Kinks from Dynamical Systems: Domain Walls in a Deformed O (N) Linear Sigma Model

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A bstract

It is shown how a integrable mechanical system provides all the localized static solutions of a deform ation of the linear O (N) -sigm a model in two space-time dimensions. The proof is based on the Hamilton-Jacobi separability of the mechanical analogue system that follows when time-independent eld congurations are being considered. In particular, we describe the properties of the dierent kinds of kinks in such a way that a hierarchical structure of solitary wave manifolds emerges for distinct N.

1 Introduction

We divide this Introduction into three parts: A.A brief history and the \state of the art". B.New developments and results to be presented in this work. C. Scenarios of possible physical applications.

Α.

K inks are solitary (non-dispersive) waves arising in several one-dimensional physical systems. Here, we shall focus on the relativistic theory of N-interacting scalar elds built on a space-time that is the (1+1)-dimensional M inkowski space R $^{1;1}$. In this context, kinks are nite energy solutions of the Euler-Lagrange equations, such that the time-dependence is dictated by the Lorentz invariance: $^{\sim}_{\rm K}$ (x;t) = $q_{\rm K}$ $\frac{{\rm r}^{\rm X}-{\rm v}^{\rm t}}{1-{\rm v}^2}$. Thus, the search for kinks leads to the solving of a system of N-coupled non-linear ordinary dierential equations and therefore becomes a very interesting problem in M athematical Physics.

The study of topological defects began as an area of research in eld theory by the mid-seventies; see [1]. It was im mediately recognized that defects of the kink type are in one-to-one correspondence with the separatrix trajectories between the bounded and unbounded motion of a mechanical system, for which the motion equations precisely form the non-linear system of dierential equations mentioned above. The equivalent mechanical system is also Lagrangian and thus automatically integrable if N = 1. For N = 2, complete integrability

is generically non-guaranteed and the equivalence to a mechanical system is not useful. This circum stance has been emphasized by Rajaraman; see [2] pp. 23-24, and partially circum vented by him self: the trial orbit method allows one to guesstimate particular types of kink trajectories.

There are, nevertheless, theories with N=2-coupled scalar elds such that the equivalent dynam ical system is completely integrable. The prototype of this kind of system is the M STB model: in [3] this was proposed in the context of the search for non-topological solitons with stability provided by a U (1) internal symmetry. In Reference [4], the model was considered as a classical continuum approximation to a 1D crystal with a two-component order parameter and it was shown that the search for kinks in this system requires that a completely integrable dynamical system be addressed. Ito, in a seminal paper [5], showed the Hamilton-Jacobi separability of the system of non-linear dierential equations. He found all the kink trajectories and explained a very peculiar kink energy sum rule.

Very rich m anifolds of kinks were discovered in two N=2 eld theoretical models, close relatives to the MSTB system, in a recent research performed by the authors of the present work [6]. The investigation of kink properties in these models requires the analysis of the separatrix trajectories in two related dynamical systems which are type I and III respectively in the classication of Liouville bidimensional (N=2) completely integrable systems, see [7].

In fact, on choosing between the four types of Liouville dynam ical systems those that m eet appropriate critical point structure, one builds an enorm ous list of related N = 2 eld theoretical models exhibiting manifolds of kinks of growing complexity; see [8]. The rôle of these models can be understood by noticing that the MSTB system is a deformation of the O(2)-linear sigma model. Instead of spontaneous symmetry breaking of O(2) by a degenerated S¹ vacuum manifold, the O(2) symmetry group is explicitly broken to Z_2 Z_2 by a mass term; only invariance under a! (1) ab a, for b = 1;2, survives. From the point of view of quantumed eld theory, this deformation is very natural because in (1+1)-dimensions infrared divergences forbid the existence of Goldstone bosons, according to a theorem of Coleman [9]. Even if it is absent in the classical action, a mass term will be generated by quantum corrections.

We interpret this as follows: in the parameter space of the N = 2 relativistic scalar eld theories invariant under the Z_2 Z_2 with generators mentioned above, and potential energy of the form

$$U(^{\sim}) = \frac{1}{2} \quad {}_{1} \quad {}_{1}^{2} + \quad {}_{2} \quad {}_{2}^{2} + \frac{1}{2} \quad {}_{1}^{4} + \quad {}_{12} \quad {}_{1}^{2} \quad {}_{2}^{2} + \frac{2}{2} \quad {}_{2}^{4} + C$$

there are at least two distinguished points. There is a choice of coupling constants such that there is explicit O (2) sym metry, which is spontaneously broken. This is the linear O (2)-sigm a model. The other interesting point is the MSTB model where the explicit Z_2 Z_2 sym metry generated by $_a$! (1) $_a$ $_b$ $_a$, for b=1;2, breaks spontaneously to the Z_2 subgroup generated by $_2$! $_2$. The key observation is that the renormalization group ow induced by quantum corrections in the parameter space avoids the O (2)-sigm a system and instead leads to the MSTB model, which also overs a variety of kinks. All the other eld theoretical models exhibiting an abundant supply of kinks also correspond to deformations of O (2)-sym metric systems with potential energies that depend on higher powers of $_1$ and $_2$, [6], [8].

There are strong analogies with the Zam olodchikov c-theorem, [10]: deform ations in the space of (1+1)-dimensional eld theories leading from conformal to integrable systems are the most interesting ones. We meet an analogous nite dimensional situation: replace the (in nite dimensional) conformal group by the O(2) group and integrability of one system with in nite degrees of freedom by integrability of a bidimensional mechanical system.

В.

This paper is devoted to investigating the kink solitary waves of the deform ation of the linear O (N)-sigm a model that generalize the M STB system to the case of N -interacting scalar elds. Non-linear waves in relativistic eld theories with N $\,$ 3 scalar elds were sketchily described for the $\,$ rst time in Reference [11]. In this work, we o er a detailed analysis of this issue. The following points meritem phasis:

- a) The dynam ical system that encodes the solitary waves of the model as separatrix trajectories has N $\,$ rst integrals in involution and hence is completely integrable. Passing from Cartesian to Jacobi elliptic coordinates in the \internal" space, R $^{\rm N}$, the dynam ical system becomes H am ilton-Jacobi separable. All the kink trajectories, and hence all the solitary waves, are then found by a special choice of the separation constants.
- b) Deep insight into the structure of the kink manifold is gained by focusing on the N = 3 case. There are three kinds of kinks: 1. A two-param eter fam ily of topological kinks with three non-null components that are \generic", i.e. they are not xed under the action of the Z_2 Z_2 Z_2 group generated by a! (1) ab a, for b = 1;2;3. 2. Four one-param eter families of \enveloping" non-topological kinks, also with three non-null components. The four families are related through the action of one Z_2 Z_2 sub-group and, together, form the envelop of the separatrix trajectories. 3. All the solitary waves of the N = 2 M STB model appear \embedded" twice; once in each plane containing the two ground states. Dierent Z_2 sub-groups leave these embedded kinks invariant.
- c) The structure of the kink manifold of the O (N) system with both explicit and spontaneous symmetry breaking repeats the patterns shown in the N = 2 (M STB) model and its generalization for N = 3. There are also generic, enveloping and embedded kinks, although when N increases the complexity of the kink manifold also increases. For instance, the N 1 kink manifold is embedded N 1 times in the manifold of kinks of the deformed linear O (N)-sigma model.
- d) In a remarkable system obtained from the generalized M STB model by also allowing asymmetries in the quartic terms of the potential, only the embedded and enveloping topological kinks living on singular edges survive as solitary wave solutions. In this system, proposed in Reference [12] for the N=2 case, the energy of all the above topological kinks is exactly the same. Together with vacuum degeneration, there is therefore kink degeneration, a phenomenon that deserves further analysis.

С.

Solitary waves of the kind that we are to describe play an important rôle in condensed matter physics. Phase transitions characterized through order parameters of the vector type

are understood in term s of the linear (or non linear) 0 (N)-sigm a m odel. The order param eter is organized in the fundam ental representation of 0 (N) and the system becomes non-linear when this N-vector is forced to take its values in the coset space M=0 (N)=0 (N 1). In (1+1)-dimensional space-time, kinks are accompanied by the fermion fractionization phenomenon [13]; this describes the continuous approximation to the bizarre behaviour of certain one-dimensional polymers such as poly-acetilene. When the spatial dimension is 3, as in the real world, kinks become domain walls which are thus related to theories involving spontaneous breaking of discrete symmetries. This happens in the hot B ig B and cosmology, where domain wall topological defects can be formed in a phase transition occurring in the expansion of the very early Universe; see [14]. More recently, domain walls have been characterized as BPS states of SUSY gluodynamics and the Wess-Zuminomodel, [15]. In all these cases there are sets of scalar elds, as in our system, that presents a variety of domain walls with dierent characteristics when seen from a 3-dimensional perspective.

In quantum eld theory, the linear O (N)-sigm a model describes systems with spontaneous symmetry breakdown to an O (N 1) sub-group and N 1 G oldstone bosons in the particle spectrum. At the beginning of the sixties G ell-M ann and Levy analyzed low energy hadronic phenomenology by introducing an elective Lagrangian eld theory of this type [16]. Besides becoming the central element of current algebra, linear sigma models also enter fundamental physics in the Higgs sector of gauge theories for elementary particle physics, see report [17] for a comprehensive review (of the perturbative sub-sector). For instance, the linear O (4)-sigma model corresponds to the Higgs sector of the electro-weak theory, while the O (24) O (5) case provides the bosonic sector of the SU (5) G rand Uni ed Theory.

Either considered on their own or forming part of G auge theories, there are reasons to discuss deformations of the linear sigmam odel. In the phenomenological approach, pions are identified with the Goldstone bosons of the model; a deformation is then necessary to convert these massless excitations in pseudo-Goldstone particles, accounting for the pions lightmass. Gauge theories are today found in the low energy limit of (fundamental) string theory. Even though deformations in the bosonic sector of gauge theories produced by small mass terms spoil renormalizability, the low energy features remain (almost) untouched and it is (almost) legitimate to trust them.

Here, we shall search for domain walls when these mild deform ations are performed in the linear O (N)-sigmam odel. It is precisely in this kind of model where the cosmological problem of wall domination is avoided [18]. Moreover, the system has a rich manifold of topological and non-topological solitons, allowing for topological defects with \internal" structure and leading to the existence of defects inside defects, a situation that generalizes a proposal of Morris [19].

The organization of the paper is as follows: In Section x2 we discuss the particle spectrum of the deformed linear O (N)-sigma model as well as the manifold of the solitary wave solutions of the system. Section x3 is devoted to the N = 3 case, which is described in full detail. We describe the situation of the generalized MSTB model for any N in Section x4 and brie y discuss the phenomenon of kink degeneration in the Bazeia system. Finally, some conclusions are drawn and some new prospects opened in Section x5. An appendix on elliptic Jacobi coordinates is also o ered.

2 Kinks in the deformed linear O(N)-sigma model

In a generic sense we understand $\$ as the solitary waves of a relativistic (1+1)-dim ensional scalar eld theory. We shall stick to the standard de nition of solitary waves; see [2] and [6]:

A solitary wave is a non-singular solution of the non-linear coupled eld equations of nite energy such that their energy density has a space-time dependence of the form:

$$"(x;t) = "(x vt)$$

where w is some velocity vector.

Given one N -com ponent scalar eld, which is a map from the R $^{1;1}$ M inkowski space-time to R $^{\rm N}$, ~ (x;t) ($_1$ (x;t); $_2$ (x;t);:::; $_{\rm N}$ (x;t)), the dynam ics of the system is governed by the action:

 $S = {}^{Z} d^{2}y \frac{1}{2} @ \sim @ \sim V (\sim)$

Here, = 0;1 are indices in the space-time and we shall use a = 1;2;:::;N to label x^N components of the eld in the \internal" R N space in such a way that \sim \sim = 1 a a . In R 1;1 we choose the metric as g = $\begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix}$ and the Einstein convention will be used throughout the paper only for the indices in R 1;1. The potential energy density is:

V (1;:::; N) =
$$\frac{2}{4}$$
 ~ ~ $\frac{m^2}{2}$ + $\frac{x^N}{4}$ $\frac{2}{4}$ a

where ;m and $_{a}$ are coupling constants of inverse length. The linear O (N)-sigm a model corresponds to the case $_{a}$ = 0,8a, which exhibits maximum O (N) sym metry. We shall focus on the deform ation of this system, which is maximally non-isotropic in the harmonic terms, i.e. $_{a}$ \in $_{b}$,8a \in b. Somehow, the deformation is natural from a quantum eld theoretical vantage point as we shall explain later and, moreover, we shall stick to the range $_{a}^{2}$ < m $_{a}^{2}$,8a, in the parameter space because in this regime the structure of the kink manifold is richer.

Introducing non-dimensional variables $! \frac{m}{}, y ! \frac{p}{m} x$ and $\frac{a}{m^2} ! \frac{2}{a}$, we not our expression for the action to be:

$$S = \frac{m^{2}}{2} d^{2}x \frac{1}{2} e^{-x} e^{-x} V(_{1}; :::;_{N})$$

$$V(_{1}; :::;_{N}) = \frac{1}{2} - _{1}^{2} + _{1}^{2} \frac{1}{2} e^{-2} e^{-2}$$

$$(1)$$

2.1 Con guration space and particle spectrum

The Cauchy problem for the eld equations

$$2_{a} = \frac{@V}{@_{a}}; \quad a = 1;2;:::;N$$
 (2)

is xed by choosing a \point" $^{\sim}$ (x;t_0) 2 M aps(R;R $^{\rm N}$) in the con guration space C, and its \tangent", $^{\sim}$ (x;t_0) 2 T_M aps(R;R $^{\rm N}$), as initial conditions to solve the system (2) of non-linear PDE.

The con guration space itself is isomorphic to the space of nite energy static con gurations; if (x;t) = q(x), C is the set of continuous maps $q:R ! R^N (q (q_1; :::;q_N))$ such that the (static) energy is nite:

$$E = \frac{m^3}{2^{12}} \frac{Z}{2} dx \frac{1}{2} \frac{dq}{dx} \frac{dq}{dx} + V(q) < +1;$$
 (3)

thus, C = fq(x) = E[q] < +1 g. $q(x) \ge C$ only if q satis es the asymptotic conditions:

$$\lim_{x! \to 1} \frac{dq_a}{dx} = 0 ; \qquad \lim_{x! \to 1} q_a(x) = v_a; 8a = 1; :::; N$$
 (4)

where \forall $(v_1; :::; v_N)$ is a constant vector that belongs to the set M of vectors annihilating V . W e assume, without loss of generality, the following ordering in the space of parameters: $_1 = 0 < _2 < ::: < _N < 1$. M is thus formed by two vectors

which are the absolute m in im a of V.

We refer to M as the vacuum manifold because in the quantum version of the theory points in M are the expectation values of the quantum eld operators \hat{a} at the ground states (\vacua") of the system . The vacuum degeneration – i.e. the existence of m ore than one vector in M – is related to the breaking of sym m etry. Besides two-dimensional Poincare invariance, there is a \internal" sym m etry with respect to the discrete group $G = Z_2$ \hat{M} :: $Z_2 = Z_2^N$ generated by \hat{a} ! (1) \hat{a} \hat{b} \hat{a} , for \hat{b} = 1;2;:::;N, 8a = 1;:::;N. The vacuum manifold is the orbit of one element by the group action

$$M = G = H_{xx} = Z_2; \qquad H_{xx} = V$$

 $H_{\,\Psi}=Z_{\,2}$ N::: $Z_{\,2}$ is the little group of the vacuum $\,\Psi$. The generators of $H_{\,\Psi}$ are the transform ations $_{a}$! (1) $_{ab}$ $_{a}$, for $b=1;2;\ldots;N$, and $a=2;3;\ldots;N$, so that $H_{\,\Psi}$ survives as a sym m etry group when quantizing around $\,\Psi$. We can understand the internal parity group G as the discrete \gauge" sym m etry: in (1+1)-dim ensions no dynam icaldegrees of freedom related to gauge potentials appear.

Vectors in M are critical points of V satisfying $\frac{\partial V}{\partial a} = 0$ and therefore constant solutions of the eld equations (2). The plane wave expansion around (x;t) = v

$$v_a$$
 (x;t) = v_a + x_a A (k) $e^{i!t ikx}$

is a solution of (2) if the dispersion relation

$$_{ab}!^{2} = _{ab}k^{2} + M_{ab}^{2}(v); M_{ab}^{2} = \frac{e^{2}V}{e_{ab}e_{b}}(v)$$
 (6)

holds.

In the quantum theory, these plane waves become the fundamental quanta with mass matrix M $_{ab}^2$ (v) and one reads the particle spectrum at a chosen critical point of V from (6). Because v 2 M are minima of V there are no negative eigenvalues of M $_{ab}^2$ (v) and the dependence on time of the plane waves around v is bounded: $e^{i!t}$. The choice of v as the starting point of the quantization procedure \spontaneously" breaks the symmetry $G = Z_2^N$ of the action to $H_v = Z_2^{(N-1)}$, which is the remaining one that survives in the particle spectrum.

In our model, we read the particle spectrum from:

$$M^{2}(\mathbf{V}) = \frac{\mathbf{m}^{2}}{2} \stackrel{\text{B}}{\underset{\text{c}}{\text{B}}} 0 \stackrel{2}{\underset{\text{c}}{\text{2}}} \cdots \stackrel{1}{\underset{\text{c}}{\text{C}}} 0 \stackrel{\text{C}}{\underset{\text{c}}{\text{C}}}$$

$$0 \quad 0 \quad \cdots \quad \stackrel{\text{C}}{\underset{\text{c}}{\text{C}}} \stackrel{\text{C}}{\underset{\text{c}}{\text{C}}}$$

$$(7)$$

Considering this system as a physical description of the continuum approximation to a one-dimensional crystal with an N-component order parameter, the particle spectrum describes a single phase with N phonon branches. We see explicitly how the symmetry group $G = Z_2^N$ is \broken" by the choice of the V vacuum to the $H_V = Z_2^{(N-1)}$ subgroup: the N phonon branches have dierent masses or \energy gaps". From the point of view of particle physics we can say that there are no tachyons; only a pseudo-Goldstone particle becomes a Goldstone boson if the corresponding a goes to zero.

It is interesting to see the model as a member of the family characterized by the potential energy densities:

$$V = \frac{1}{2} {\binom{2}{3}} {\binom{N^{N}}{2}}$$
ab a b
$${}^{2A} + {\binom{N^{N}}{2}} {\binom{ab}{2}}$$
a b

where $_{ab}$; $_{ab}$ and 2 are \bare" non-dim ensional param eters. Ultraviolet divergences are controlled by normal ordering in the quantum theory, but the need arises to introduce a renormalization 'point' 2 , and the dependence of the renormalized parameters on 2 is determined by the renormalization group equation. One special solution, a special renormalization group ow, might lead to the \point":

$$_{ab}^{R}(^{2}) = _{ab}; \quad _{ab}^{R}(^{2}) = 0; \quad ^{R}(^{2}) = 1$$

in the space of quantum eld theory models in the family. This point is the linear O (N)-sigm a model which has G=0 (N) as the (continuous) sym metry group. The vacuum orbit is, however, M=0 (N)=0 (N 1)= S^{N-1} , the (N 1)-dimensional sphere, and thus there is no unbroken sym metry left: there are N 1 massless particles. If the only modication of the renormalized parameters is to allow for non-zero values of $\frac{R}{ab}(^2)$; a=b=r+1;:::;N there are still r 1 G oldstone bosons.

Colem an [9] established that in (1+1)-dimensions the infrared asymptotics of the two-point Green functions of a quantum scalar eld forbids poles at $!^2 = k^2$; there are no Goldstone bosons in (1+1)-dimensions. It is thus impossible to reach the O (N)-sigma model or its deformation with the O (r) symmetry spontaneously broken to O (r 1) in the renormalization group ow. The closest admissible points are the models characterized by:

$$_{ab}^{R}(^{2}) = _{ab};$$
 $^{R}(^{2}) = 1;$ $_{ab}^{R} = 0;$ a $\in b$

$${R \choose 11}({}^2) = 0 < {R \choose 22}({}^2) = {2 \choose 2} ::: {R \choose NN}({}^2) = {2 \choose N} < 1$$

In this paper we shall focus on the case of maximal explicit symmetry breaking; i.e. when strict inequalities in the parameter space occur. Nevertheless, we shall comment on the allowed situation characterized by

$$\frac{2}{1} = 0 < \frac{2}{2} = :::= \frac{2}{r_1} < \frac{2}{r_1+1} = :::= \frac{2}{r_2} < :::< \frac{2}{r_2+1} = :::= \frac{2}{N} < 1$$

when there is degeneration in the spectrum but no G oldstone bosons. Note that the generators of the O $(r_1 \ 1) \ O \ (r_2 \ r_1) \ \dots \ O \ (N \ r_k)$ sym m etry sub-group are in the little group of the vacuum . The sym m etry group is $G = Z_2 \ O \ (r_1 \ 1) \ O \ (r_2 \ r_1) \ \dots \ O \ (N \ r_k)$, $H_{\forall} = O \ (r_1 \ 1) \ O \ (r_2 \ r_1) \ \dots \ O \ (N \ r_k)$ and the vacuum orbit is $M = G = H_{\forall} = Z_2$.

2.2 Con guration space topology: kinks and dynam ical system s

The con guration space of the model is the union of topologically disconnected sectors:

$$C = {\begin{picture}(20,0) \put(0,0){\line(1,0){100}} \put(0,0){\line(1,0$$

hom otopy group of C and its order. This comes from the asymptotic conditions (4) and the continuity of the time evolution. There are topological charges dened for each conguration in C as: $\[\]$

$$Q_a^T = \frac{1}{2}^{\frac{Z}{1}} dx \frac{d_a}{dx} = \frac{1}{2} (_a(+1;t))$$

It should be noted that Q_a^T is independent of t, 8a, and in our system equal to zero if a 2. Therefore the four sectors C are labelled by the values; of the elds at in nity compatible with nite energy and Q_1^T determines the homotopy class in Q_1^T (C) = Z_2 Z_2 .

The critical points of E are time-independent nite-energy solutions of the eldequations. If they are not spatially hom ogeneous, the critical points correspond to solitary waves that are therefore related to the topological structure of C. Besides complying with (4), solitary waves satisfy the system of ordinary dierential equations:

$$\frac{\mathrm{d}^2 q_a}{\mathrm{d}x^2} = \frac{\mathrm{eV}}{\mathrm{eq}_a} \tag{8}$$

Recall that $_a(x;t)=q_a(x)$. Solving the system (8) is tantam ount to noting the solutions of the Lagrangian dynamical system in which x=p plays the rôle of time, the \particle" position is determined by $q_a()$, and the potential energy of the particle is U(q)=V(q). From this perspective the static eldenergy E is seen as the particle action:

$$E = J = \begin{bmatrix} Z & (1) \\ \frac{1}{2} \frac{dq}{d} & \frac{dq}{d} \end{bmatrix} U (q)$$
 (9)

Trajectories that behaves asym ptotically in the —time as ruled by (4) have a nite action, J, in them echanical problem and are in one-to-one correspondence with solitary waves/kinks that have energy E=J in the eld theoretical system.

The mechanical analogy is very helpful when one is dealing with a real scalar eld theory because, then, a rst integral is all that we need to nd all the solutions. Vector scalar

elds of N components lead to N -dimensional dynamical systems which are seldom solvable. Magyari and Thomas [4] realized that the two-dimensional dynamical system arising in connection with the MSTB model is a completely integrable one in the Liouville sense; there are two rst integrals in involution. Moreover, Ito [5] has shown that the mechanical system is Hamilton-Jacobi separable, noting all the trajectories and hence all the kinks of the MSTB model. In a recent publication [6], we have developed this procedure for two N=2 models with interesting features: the rst system is a deformation of the (1+1)-dimensional scalar eld theory, where the potential energy density is the Chem-Simons-Higgs potential arising in self-dual planar gauge theories. The second one is a deformation of the linear O (2)-sigma model, which is dierent from the MSTB model.

To extend thism ethod of nding kinks to the linear O (N)-sigm a m odel, N = 3, deform ed in such a way that the O (N) sym m etry is explicitly broken to G $= \mathbb{Z}_2^N$, we start from the \particle" action:

$$J = {\overset{Z}{d}} {\overset{(1)}{d}} {\overset$$

The particle motion equations are:

$$\frac{d^2q_a}{d^2} = 2q_a(q q 1) + {}^2_aq_a; 8a = 1; :::; N$$
 (10)

which are mathematically identical to the eld equations for static congurations. Finite action trajectories, kinks in the eld theory, should also satisfy the asymptotic conditions:

$$\lim_{!} \frac{dq_{a}}{d} = 0; \qquad \lim_{!} q_{a}() = a_{1}$$
 (11)

We shall use the H am iltonian form alism to integrate the mechanical system. The canonical momenta p_a () = $\frac{\partial L}{\partial q_a}$ = $\frac{dq_a}{d}$ (), together with the positions q_i (), form a system of local coordinates in phase space. We should bear in mind that p_a () = $\frac{d}{dx}$ when going back to the eld theory. The mechanical H am iltonian

$$I_1 = \frac{1}{2}p \quad p \quad \frac{1}{2} (q \quad q \quad 1^2) \quad \prod_{a=1}^{x^N} \frac{1}{2} \quad {}_a^2 q_a^2$$
 (12)

leads to the system of canonical equations

$$\frac{dq_a}{d} = fI_1; q_a g; \qquad \frac{dp_a}{d} = fI_1; p_a g$$

equivalent to (10). Given any two functions F(q;p), G(q;p) in phase space, the Poisson bracket is de ned in the usual way:

O by iously $\frac{dI_1}{d} = 0$, but our mechanical system is full of other invariants. In fact, as early as 1919 G amier [20] solved the motion equations and described periodic trajectories in terms

of Theta functions: the kink trajectories of nite \action" correspond to a limiting case and are the separatrices between the periodic trajectories and unbounded motion. More recently Grosse, and other authors [21] have shown that the functions:

$$K_{a} = \sum_{b=1; b \in a}^{X^{N}} \frac{1}{\sum_{a=2}^{2} 1_{ab}^{2} + p_{a}^{2} + (2 - \frac{2}{a})q_{a}^{2} - q_{a}^{2} + q_{b}^{2}}$$

$$I_{ab} = p_{a}q_{b} - p_{b}q_{a}$$

$$(13)$$

are rst integrals in involution:

$$fI_1 ; K_a q = 0$$
 $fK_a ; K_b q = 0$

There is a set of N + 1 invariants in involution: I_1 ; K_1 ; K_2 ; I_1 ; I_2 ; I_3 ; I_4 ; I_5 ; I_7 ; I_8 ;

At this point we pause to explain the singular nature of the deform ation of the linear O (N)-sigm a model chosen from among many possibilities. The $\frac{2}{a}$ term s explicitly break the O (N)-sym metry of the linear sigm a model; the case $\frac{2}{a}=0$, 8a = 1;2;:::;N . In the mechanical system the O (N) internal transform ations become ordinary rotations. The angular momentum components, l_{ab} , conserved in the lim it $\frac{2}{a}=0$, 8a, are no longer 'time'-independent if $\frac{2}{a}$ 6 0. There are, however, N invariants K a, which in the lim it a=0, 8a, are given in terms of the O (N)-invariants: the r C asim ir invariants and the r generators of the C artan sub-algebra, where $r=\frac{N}{2}$ or $\frac{N-1}{2}$ if N -even or -odd is the rank of the group. A warning: in the N = odd case, the energy must be added to the other N = 1 invariants built from the C artan sub-algebra and the C asim ir invariants. For any N , the maximally asymmetric chosen deformation is special because it retains enough symmetry to solve the mechanical system. There is no Lie algebra associated with K a however; since the invariants are quadratic in q_a , p_a , the action of K a in the phase space, given by fK a; $q_b g$ and fK a; $p_b g$, is non-linear.

In (1+1)-dim ensional eld theory, the energy-m om entum tensor:

$$T = \frac{@L}{@(@_a)} \quad @_a \quad g \quad L$$

is divergenceless due to invariance under space-time translations. $P = {R \atop a} dx T^0$ are thus conserved quantities whatever the values of ${A \atop a}$. The O (N) \isospin" currents however,

$$J_{a} = X^{N} \quad C_{abc b} 0 \quad c$$

are only divergenceless if $_a$ = 0,8a. The c_{abc} are the Lie 0 (N) structure constants and the charges Q_a = R dx J_a^0 are not conserved if there is no symmetry with respect to the transform ation generated by them . For static con gurations, we have

$$dxT^{00} = E = J; T^{10} = T^{01} = 0; T^{11} = I_1$$

$$J_a^0 = 0;$$
 $J_a^1 = X^N c_{abc} l_{bc}$

In term s of the 'isospin' currents, the invariants K a can be written as:

We expect that the time-evolution occurs in such a way that there is some equation of non-linear character!

$$F = \frac{\partial L_a}{\partial t}; \frac{\partial K_a}{\partial x} = 0$$

between K_a and

which reduces to $\frac{\theta J_0^a}{\theta t} = \frac{\theta J_0^a}{\theta x}$ when a = 0, 8a. The situation is analogous to that occurring between conformal eld theories and models with in nite-dimensional algebraic symmetry as in (1+1)-dimensional Toda eld theories and Toda a nemodels [22]. There are two dierences: (1) the conformal group is in nite dimensional in (1+1)-dimensions. We have only one nite-dimensional group O (N) and thus we can solve only the static limit of the eld theory model. (2) Due to the non-linear character of the deformation of the O (N) Lie generators, we do not even have a nite-dimensional Lie algebra.

2.3 The Hamilton-Jacobi equation and kink trajectories

The K $_a$ invariants de ned in (13) are quadratic in the momenta, but not orthogonal (they contain terms in p_ap_b ; $a \in b$). Therefore, the Stackel theorem can not be applied to assure H am ilton-Jacobi separability. This problem is surpassed in our dynamical system with the choice of some suitable system of coordinates. The appropriate system is provided by elliptic Jacobi coordinates, with a choice of separation constants determined by the deformation parameters giving mass to the Goldstone bosons; $_a^2 = 1$ $_a^2$; 8a = 1;2;:::;N . Thus we dene:

$$q_{a}^{2} = \frac{\frac{Y}{A}}{\frac{b=1}{Y}} \left(\begin{array}{cc} 2 & b \\ a & b \end{array} \right) = \frac{\left(\begin{array}{cc} 2 \\ a \end{array} \right)}{A^{0} \left(\begin{array}{cc} 2 \\ a \end{array} \right)}$$

$$b=1; b \in a$$
(14)

ruling the change of coordinates from Cartesian, q (q_1 ;:::; q_N) to elliptic \sim ($_1$;:::; $_N$). In the Appendix, it is explained how the elliptic variables are split:

$$1 < {}_{1} < {}_{N} < {}_{2} < {}_{N} {}_{1} < ::: < {}_{2} < {}_{N} < 1$$
 (15)

Notice that formula (14) coincides with formula (72) in the Appendix if we change q_a by q_{N-a+1} and choose $r_{N-a+1} = \frac{2}{a}$.

Together with form ula (14), this splitting m eans that the change of coordinates produces a map from a sub-space of R $^{\rm N}$, characterized as the set of points which are not invariants under the Z $_2^{\rm N}$ group generated by q $_1$! (1) $^{\rm ab}q_a$;b = 1; ;N , to the interior of the in nite parallelepiped P $_{\rm N}$ (1) obtained by replacing the inequalities in (15) by equalities: 1 < $_1$ $_{\rm N}^2$::: $_2^2$ $_{\rm N}$ 1. Notice that in this map $2^{\rm N}$ regular points in R $^{\rm N}$ go to a single point in the interior of P $_{\rm N}$ (1); Singular points lie in the R $^{\rm m}$, m = 0;1;:::;N 1, sub-spaces that are invariant under the action of some non-trivial element of G = Z $^{\rm N}$. These singular sub-spaces are mapped into the boundary of P $_{\rm N}$ (1).

The standard length of an interval in Euclidean space is expressed in elliptic coordinates in the form

$$ds^2 = \int_{a=1}^{x^N} dq_a dq_a = \int_{a=1}^{x^N} g_{aa} (^{\sim}) d_a d_a$$

because the m etric g_{aa} ($^{\sim}$), as derived in the Appendix, is:

$$g_{aa}(^{\sim}) = \frac{1}{4} \frac{f_a(^{\sim})}{A(_a)}; \quad g_{ab}(^{\sim}) = 0;8a \in b$$

where A (
$$_{a}$$
) = $_{b=1}^{W}$ ($_{a}$ $_{b}^{2}$), and 1

$$f_a(^{\sim}) = f_a(_1;_2;_{N};_{D_{ba}}) = \int_{b=1}^{N} (_a _b)$$

Therefore, the Lagrangian reads:

$$L = \frac{1}{2} \sum_{a=1}^{X^{N}} \frac{dq_{a}!}{d} U(q_{1}; :::; q_{N})$$

$$= \frac{1}{2} \sum_{a=1}^{X^{N}} g_{aa}(^{\sim}) \frac{d_{a}!}{d} U(q_{1}; :::; q_{N})$$
(16)

where the potential in elliptic coordinates is:

$$U (_{1};:::;_{N}) = \frac{x^{N}}{a=1} \frac{1}{2} \frac{\frac{N+1}{a} (_{1}) \frac{N}{a} + (1 + _{1}) \frac{N}{a} \frac{1}{a}}{f_{a}(^{\sim})}$$

$$= \frac{x^{N}}{a};_{a=1} = \frac{x^{N}}{a};_{a=1} = \frac{x^{N}}{a} \frac{x^{N}}{a}$$

$$= \frac{x^{N}}{a};_{a=1} = \frac{x^{N}}{a} \frac{x^{N}}{a}$$

The computation of U ($^{\sim}$) is highly non-trivial and requires the use of form ulas that follow the Jacobi Lem ma, such as (76), (77), (78), etc.

Appendix. We shall use f_a (~) in the main text instead of $\,^0$ ($_a$), to stress the fact that this quantity depends on all the components of ~

¹The standard notation in the literature on elliptic Jacobi coordinates is $^{0}(_{a})=_{_{b=1}^{b=1}}^{Y^{N}}(_{a})$, see

The canonical momenta associated to the a variables are:

$$a = \sum_{b=1}^{N} g_{ab}(^{\sim}) \frac{d_{b}}{d} = g_{aa}(^{\sim}) \frac{d_{a}}{d}$$

and, through the standard Legendre transform ation, we write the Hamiltonian:

$$H = \frac{1}{2} \frac{x^{N}}{a_{a=1}} \frac{4A(a)}{f_{a}(a)} a^{2} + U (a)$$
 (18)

The key point is that H can be written in Stackel's form:

$$H = \frac{x^{N}}{a=1} \frac{H_{a}}{f_{a}(\tilde{a})} = (19)$$

$$= \frac{x^{N}}{a=1} \frac{2A(a)^{\frac{2}{a}} \frac{1}{2} x^{N+1} (1)^{\frac{N}{a}} + (1+1)^{\frac{N}{a}}}{f_{a}(\tilde{a})}$$

such that the Hamilton-Jacobi equation

$$\frac{\text{@S}}{\text{@}} + H \quad \frac{\text{@S}}{\text{@}_{1}}; \dots; \frac{\text{@S}}{\text{@}_{N}}; _{1}; \dots; _{N} = 0$$
 (20)

is completely separable. We now prove this last statement.

Fixing $H = I_1$, the rst integral of energy, and having in m ind the expression (19) of H, we write the solution of (20) as:

$$S = I_1 + \sum_{a=1}^{\hat{X}^N} S_a(a):$$
 (21)

Therefore, (20) reduces to:

$$I_{1} = H \quad \frac{dS_{1}}{d_{1}}; \dots; \frac{dS_{N}}{d_{N}}; _{1}; \dots; _{N} = \frac{X^{N}}{d_{n}} \frac{H_{a}}{f_{a}(^{\sim})}$$
(22)

The Ham ilton-JacobiPDE equation (20) becomes equivalent to the system of non-coupled ordinary dierential equations

$$H_{a} \frac{dS_{a}}{d_{a}}; a = {}_{1} {}_{a} {}^{N} {}_{1} + {}_{2} {}_{a} {}^{N} {}_{2} + :::+ {}_{N} {}_{1} {}_{a} + {}_{N}$$
 (23)

w here

$$H_{a} \frac{dS_{a}}{dz}; a = 2A(a) \frac{dS_{a}}{dz} = \frac{1}{2} \sum_{a=1}^{N+1} (1)^{a} + (1 + 1)^{a} ; (24)$$

due to the identity

$$I_{1} = I_{1} \frac{x^{N}}{a=1} \frac{x^{N}}{f_{a}(^{\sim})} + 2 \frac{x^{N}}{a=1} \frac{x^{N}}{f_{a}(^{\sim})} + \dots + x^{N} \frac{1}{f_{a}(^{\sim})}$$
(25)

-observe that $_1 = I_1 - w \, hich$ follows from the Jacobi Lemma and the subsequent relations (Appendix)

$$\frac{X^{N}}{a} = \frac{1}{f_{a}(^{\sim})} = 1;$$
 $\frac{X^{N}}{a} = \frac{1}{f_{a}(^{\sim})} = 0;8i = 2;...;N$

A liternatively, one could take a more direct approach to show formula (23). We start from (22), written explicitly as

$$\frac{H_{1}(1)}{(1 2)(1 3)} + \frac{H_{2}(2)}{(2 1)(2 3)} + \frac{H_{1}(1)}{(1 2 1)(2 3)} + \frac{H_{1}(1)}{(1 2 1)(2 3)} = I_{1} (26)$$

Here, each $H_a(a)$ is of the form (24) due to the ansatz (21). Multiplying (26) by (a_2) and setting $a_1 = a_2$ one sees that $H_1(a_1) = a_2$ is a setting $a_1 = a_2$ one sees that $H_1(a_1) = a_2$ is a setting $a_1 = a_2$ in a setting $a_1 = a_2$ in a setting $a_1 = a_2$ is a setting $a_1 = a_2$ one sees that $H_1(a_1) = a_2$ is a setting $a_1 = a_2$ in a setting $a_1 = a_2$ in a setting $a_1 = a_2$ is a setting $a_1 = a_2$ in a setting $a_1 = a_2$ in a setting $a_1 = a_2$ in a setting $a_1 = a_2$ is a setting $a_1 = a_2$ in a setting $a_2 = a_1$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a setting $a_2 = a_2$ in a setting $a_1 = a_2$ in a setting $a_2 = a_2$ in a settin

$$H_1(_1) + P_2(^{\sim})H_2(_2) + \qquad \qquad \uparrow P)H_N(_N) = I_1(_1 __2)(_1 __3) \qquad \qquad _1(_N); \quad (27)$$

where each $P_a(\tilde{\ })$, a=2;3; ;N , is a polynomial of degree N 1 in. Dierentiating (27) N times with respect to 1 yields

$$\frac{d^{N} H_{1}(1)}{d^{N}} = N I_{1}$$
 (28)

whence it follows that $H_1()$ is a degree N polynomial in with leading coe cient I_1 , in agreement with equation (23).

The set of separation constants $_{i}$; i=1;2;:::;N is another system of rst integrals in involution. They can be expressed in the elliptic phase space T P_N (1) as functions of $_a$ and $_a=\frac{dS_a}{d_a}$ by solving the linear system of equations (23) in the unknown $_i$, which is a Vanderm onde system . The N $_i$ 1 roots of the polynom ial

$$\frac{N}{a}^{1} + \frac{2}{I_{1}}^{N} = \frac{N}{a}^{1} + \dots + \frac{N}{I_{1}} = (a F_{2})(a F_{3}) \dots (a F_{N})$$

together with the energy $F_1 = I_1$ form another system of invariants in involution. Both systems are related through the identities $_1 = I_1$ and:

$$_{a} = (1)^{a}I_{1}$$
 $_{i_{1} < i_{2} < ::: < i_{a}}^{X} F_{i_{1}}F_{i_{2}} ::: F_{i_{a}}; i = 2;3; ::: ;N$

Therefore, all the separation constants $_a$ are proportional to the \particle" energy I_1 . Dening the polynom ialB ($_a$) in the form :

B(a) =
$${}^{N+1}_{a}$$
 (1) N + (1 + 2₁) ${}^{N}_{a}$ 1 + 2₂ ${}^{N}_{a}$ 2 + :::+ 2_N

the solution of the di erential equation (23) is a quadrature:

$$S_{a}(a) = \frac{1}{2} \operatorname{sign} \frac{dS_{a}}{da} \stackrel{!}{=} \frac{V}{A} \frac{V}{A} \frac{B(a)}{A(a)} da \qquad (29)$$

and the general solution of the Hamilton-Jacobi equation reads:

$$S = {}_{1} + {}_{a=1}^{X^{N}} \frac{1}{2} sign \frac{dS_{a}}{da}! Z \overset{V}{u} \frac{\overline{B(a)}}{A(a)} da$$
 (30)

The explicit integration of the quadratures in (29) requires the theory of Theta functions of genus depending on N . The action of the associated trajectories is in nite because they are either periodic or unbounded. The asymptotic conditions (4) that guarantee nite action to continuous trajectories satisfying them also require that the energy used by the particle in these trajectories should be zero. This is so because $I_1j_{=1}=0$ and, being an invariant of the evolution, $I_1=0$;8 .

We recall that the trajectories of nite action in an evolution lasting an in nite time are the kinks of the parent eld theory system: one just trades the nite action of the trajectory for nite energy of the non-linear wave. Therefore, the kinks are the trajectories obtained when all the separation constants in (23) are zero: $_{a} = 0$;8a. These are the separatrices between bounded and unbounded motion and the integrals in (29) are easier to compute.

The explicit trajectories are also provided by the Hamilton-Jacobi principle, through the set of equations:

$$a = \frac{@S}{@a}$$
; $a = 1;2;...;N$

where the $_a$ are integration constants. In the hypersurface of the phase space determ ined by $_1$ = :::= $_N$ = 0, the rst equation

$$1 = + \frac{x^{N}}{2} \frac{1}{2} sign(a)^{\frac{N}{a}} \frac{1}{3^{N}(a)^{\frac{N}{a}}} \frac{1}{a^{N+1}} \frac{A(a)}{(1)^{\frac{N}{a}} + (1 +)^{\frac{N}{a}}}$$
(31)

rules the time-dependence of the particle in its journey through the orbit. From the eld theoretical point of view, it provides the kink form factor. The other N 1 equations, i=2;3;:::;N,

$$i = \sum_{a=1}^{X^{N}} \frac{1}{2} sign(a)^{Z} \frac{\sum_{a=1}^{N} id_{a}^{V}}{\sum_{a=1}^{N} (a)j} \frac{A(a)}{\sum_{a=1}^{N} (a)j} \frac{A(a)}{\sum_{a=1}^{N} (a)j}$$
(32)

determ ine the orbit in P_N (1), the intersection of N 1 hypersurfaces in the con guration space. Therefore, there is a N 1-dim ensional family of kinks parametrized by the nite values of $\frac{1}{2}$.

Although (31) and (32) identify all the separatrix trajectories of the mechanical system and henceforth all the kink solutions of the deformed linear O (N) -sigma model, an explicit description of such solitary waves is dicult for two reasons: (1). (31) and (32) form a system of transcendent equations of impossible analytical resolution. (2). Even if it were possible, expressing back the solution in Cartesian coordinates through (72) for N 3 is another in possible task by analytical means.

3 N = 3

To gain insight into the nature of the di erent kinks of the model, in this Section we shall address in full detail the N=3 case. We shall deal with a (1+1)-dimensional eld theory including three scalar elds which transform according to a vector representation of the O (3) group. The structure of the solitary wave solutions of the N=3 system is extremely rich from dierent points of view and shows the behavioural pattern of the general case with N-component elds.

3.1 The general solution of the Hamilton-Jacobi equation

The Hamiltonian of the underlying dynamical system reads:

$$H = \frac{1}{2} p_1^2 + p_2^2 + p_3^2 \frac{1}{2} q_1^2 + q_2^2 + q_3^2 \frac{1}{2} q_2^2 \frac{\frac{2}{3}}{2} q_2^2 \frac{\frac{2}{3}}{2} q_3^2$$
 (33)

in Cartesian coordinates. To write the Ham iltonian in elliptic coordinates, note that for N = 3 we have:

$$=1+ \ \ ^2_2 + \ ^2_3 \ ; \qquad = \ \ ^2_2 \ ^2_3 + \ ^2_2 + \ ^2_3$$

$$A(\ _a) = (\ _a \ \ 1)(\ _a \ \ ^2_2)(\ _a \ \ ^2_3)$$

$$^0(\ _1) = (\ _1 \ \ _2)(\ _1 \ \ _3); \ ^0(\ _2) = (\ _2 \ \ _1)(\ _2 \ \ _3); \ ^0(\ _3) = (\ _3 \ \ _1)(\ _3 \ \ _2)$$

$$H \ ence, H \ = \ \ ^{X^3} \ \ \frac{1}{^{0}(\ _a)} H_a \ , \ where$$

$$H_{a} = 2A \left(a \right)^{2} \frac{1}{a} \frac{1}{2} \frac{1}{a} \left(a \right)^{2} \left(a \right)^{2} \left(a \right)^{2}$$
(34)

The separatrix trajectories, those in one-to-one correspondence with solitary waves of kink type in the encompassing eld theory, are fully determined by the equations (32) restricted to the N=3 case:

$$C_{2} = \begin{array}{c} P_{\frac{1}{1} + 2} \\ P_{\frac{1}{1} + 2}$$

where $C_2 = \exp f2_{2} \cdot 2_3 (\frac{2}{2} \cdot \frac{2}{3}) g$ is constant, and:

$$C_{3} = \begin{array}{c} \begin{array}{c} p \\ \hline 1 \\ \hline 2 \\ \hline 1 \\ \hline 1 \\ \hline 2 \\ \hline 1 \\ \hline 1 \\ \hline 2 \\ \hline 1 \\ \hline 1 \\ \hline 2 \\ \hline 2 \\ 3 \\ 3 \\ \text{sign(1)} \end{array} \right)$$

with $C_3 = \exp f2_{3} + 2_{3} + 2_{3} + 2_{3} + 2_{3} = 2_{3} + 2_{3} = 2_{3}$

Integration of (31) in the N = 3 case shows the time-table of the particle in each trajectory, or, the kink form factor:

$$C_{1}() = \begin{array}{c} P \\ \hline 1 \\ \hline 1 \\ \hline 1 \\ \hline 1 \\ \hline 2 \\ \hline \end{array} \begin{array}{c} 3 \\ \frac{2}{2} sign(1) \\ \hline P \\ \hline \hline 1 \\ \hline 1 \\ \hline \end{array} \begin{array}{c} 1 \\ \hline 1 \\ \hline \end{array} \begin{array}{c} 1 \\ \hline \end{array} \begin{array}{c} 2 \\ \frac{2}{3} sign(1) \\ \hline \end{array} \\ P \\ \hline \hline \end{array} \begin{array}{c} 1 \\ \hline 2 \\ \hline \end{array} \begin{array}{c} 3 \\ \frac{2}{2} sign(2) \\ \hline \end{array} \begin{array}{c} P \\ \hline \hline 1 \\ \hline \end{array} \begin{array}{c} 2 \\ \frac{2}{3} sign(2) \\ \hline \end{array} \begin{array}{c} 2 \\ \frac{2}{3} sign(2) \\ \hline \end{array} \begin{array}{c} 2 \\ \frac{2}{3} sign(2) \\ \hline \end{array} \begin{array}{c} 2 \\ \frac{2}{3} sign(3) \\ \hline \end{array} \begin{array}{c} P \\ \hline \end{array} \begin{array}{c} 2 \\ \frac{2}{3} sign(3) \\ \hline \end{array}$$

if C_1 () = expf2($_1$ +)($_3^2$ $_2^2$) $_2$ $_3$ g. Therefore, there is a fam ily of kinks param etrized by the integration constants $_2$, $_3$: it corresponds to the fam ily of curves in P_3 (1) determ ined by the intersection of the surfaces de ned by (35) and (36). The third constant, $_1$, xes the center of the kink, the point where the energy density reaches its maximum value.

Better intuition of the kink shapes requires an interpretation of the solutions described by equations (35) and (36) in Cartesian coordinates. We shall describe how is this achieved in the next sub-sections, but before this it is convenient to note some details of the change of coordinates from Cartesian to elliptic in R³:

$$q_{1}^{2} = \frac{1}{\frac{2}{2} \frac{2}{3}} (1 \quad 1)(1 \quad 2)(1 \quad 3)$$

$$q_{2}^{2} = \frac{1}{\frac{2}{2} (\frac{2}{3} \quad \frac{2}{2})} (\frac{2}{2} \quad 1)(\frac{2}{2} \quad 2)(\frac{2}{2} \quad 3)$$

$$q_{3}^{2} = \frac{1}{\frac{2}{3} (\frac{2}{2} \quad \frac{2}{3})} (\frac{2}{3} \quad 1)(\frac{2}{3} \quad 2)(\frac{2}{3} \quad 3)$$
(38)

The change of coordinates is singular at the three R 2 coordinate planes; $q_1 = 0$, $q_2 = 0$ and $q_3 = 0$. The in age of the $q_1 = 0$ plane is a unique face, $q_1 = 1$, of the P $_3$ (1) parallelepiped:

$$1 < \frac{2}{3} \quad 2 \quad \frac{2}{2} \quad 3 \quad 1$$
 (39)

The $q_2=0$ plane, however, is mapped into faces $_2=\frac{2}{2}$ and $_3=\frac{2}{2}$, while the $q_3=0$ plane goes to faces $_2=\frac{2}{3}$ and $_1=\frac{2}{3}$ of $P_3(1)$. Observe that $g_{11}(\frac{2}{3};_2;_3)=g_{22}(_{1};_{2}^{2};_{3})=g_{33}(_{1};_{2};_{2}^{2})=g_{33}(_{1};_{2};_{1})=1$. The whole R 3 space

is mapped in $P_3(1)$. Due to the symmetry under the group $G=Z_2-Z_2$ generated by q_a ! q_a , the mapping (38) is eight to one in regular points of R^3 : to any point in the interior of $P_3(1)$ correspond eight points in R^3 away from the coordinate planes. These planes are xed loci of some subgroup of G.

The asymptotic conditions (4) in q_a restrict the motion to the compact sub-space D 3 of R 3 bounded by the tri-axial ellipsoid:

$$q_1^2 + \frac{q_2^2}{\frac{2}{3}} + \frac{q_3^2}{\frac{2}{3}} = 1 \tag{40}$$

and are satis ed by nite action and zero energy trajectories. Elliptic coordinates are best suited for demonstrating such a restriction. In this coordinate system D^3 is mapped to the nite parallelepiped $P_3(0)$:

$$0 \quad {}_{1} \quad {}_{3}^{2} \quad {}_{2} \quad {}_{2}^{2} \quad {}_{3} \quad 1 \tag{41}$$

The unique non-singular face of $P_3(0)$ with respect to the change of coordinates is $_1=0$ and the inverse in age of this face is the ellipsoid (40). The asymptotic conditions (4) force $_b=0$;8b, and thus, by (23), $H_a=0$;8a, for the nite action solutions. $_2$ and $_3$ are bounded, see (41). Thus, we focus on,

$$H_1 = 0$$
) $\frac{1}{2} {}_1^2 + \frac{1}{8} \frac{{}_1^2}{{}_1} = 0$ (42)

Equation (42) describes the motion of a particle with zero energy moving under the in uence of a potential

$$V(_{1}) = \frac{1}{4} \frac{_{1}^{2}}{_{1}};$$
 $1 < _{1} _{3}^{2}$
= 1; $_{3}^{2} < _{1} < 1$

V($_1$) has a maximum at $_1$ = 0 and goes to 1 when $_1$ tends to 1; therefore, bounded motion occurs only in the $_1$ 2 [0; $_3^2$] interval and the trajectories giving rise to kinks lie in P₃(0), seen in elliptic coordinates, or D³ in Cartesian space.

In Figure 1 the whole picture is depicted and we notice the following important elements of the dynamics:

-Points: (1) the origin. This is a xed point of $G=Z^3$ and thus only one point 0 in D³ is mapped to the vertex 0 in P₃(1). (2) Points B, C, D: these are the intersection points of the three distinguished ellipses , $q_1^2 + \frac{q_2^2}{\frac{2}{2}} = 1$, $q_1^2 + \frac{q_3^2}{\frac{2}{3}} = 1$ and $\frac{q_2^2}{\frac{2}{2}} + \frac{q_3^2}{\frac{2}{3}} = 1$, in the ellipsoid (40). They are xed points under the action of a sub-group Z_2^2 of G and thus, two points in the boundary of D³ are mapped to a single point in the boundary of P₃(0). D is the point where the two vacuum points v are mapped and hence it is a very in portant point of the dynamics: every nite action trajectory starts and ends at D. (3) Points F₁, F₂, F₃, the foci of the above ellipses. Again two points in D³ are mapped in a unique point in P₃(0). (4) The umbilicus A of the ellipsoid $_1 = 0$ is another characteristic point; in Cartesian coordinates A corresponds to four points in the boundary of D³ because they are invariant only under a Z_2 sub-group of G.

-Curves:

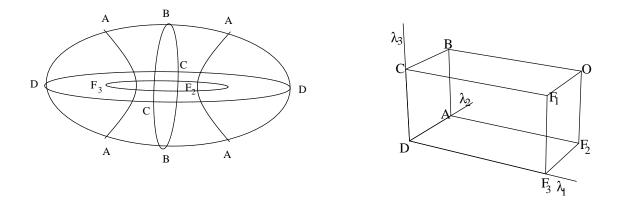


Figure 1: (a) The dom ain D 3 in R 3 : Cartesian coordinates (b) The dom ain P $_3$ (0) in R 3 : Jacobi elliptic coordinates

The ellipse with foci F_2

$$\frac{q_1^2}{\frac{2}{3}} + \frac{q_2^2}{\frac{2}{3} + \frac{2}{3}} = 1 \tag{43}$$

in the $q_3=0$ plane passing through F_3 and F_1 . This is the edge $_2=_3^2=_1$ in $P_3(0)$. Observe that four points on the ellipse (43) are mapped to one point in the edge of $P_3(0)$, because it is invariant under a Z_2 sub-group of G. The map leading to F_3 and F_1 is, however, two to one: the invariance group of these points is bigger, Z_2^2 G.

The hyperbola

$$\frac{q_1^2}{\frac{2}{3}} \quad \frac{q_3^2}{\frac{2}{3} \quad \frac{2}{3}} = 1 \tag{44}$$

in the $q_2 = 0$ plane passing through F_2 and A and having foci F_3 (the edge $_2 = _2^2 = _3$ in $P_3(0)$).

The above points and curves play a special rôle in the de nition of the elliptic coordinates and are also \critical loci" of the dynam ics.

3.2 Generic Kinks

The generic kinks of the N=3 system are the trajectories given by the solutions of (35)-(36) for non-zero nite values of C_2 and C_3 . The solutions of the implicit equations (35)-(36) cannot be graphically represented by means of the built-in functions of M athematica. We use a numerical algorithm implemented in M athematica to obtain the graphic portrait of the trajectories. The algorithm allows us to calculate an arbitrary number of points on the orbit. These points joined by straight segments provide a visualization of the trajectory. There is a special step and an iteration of routine steps in the procedure, which is based on the New ton-Raphson m ethod.

First step. Identication of two points on the trajectory.

For given values of C_2 , C_3 , C_3 , C_3 , and a choice of signs, we set the rst variable to the point" $C_1 = C_1$. (35)-(36) becomes a system of two equations in two unknowns that can be

solved by the New ton-Raphson method with starting values (0_2 ; 0_3). The outcome is a point P_1 (1_1 ; 2_2 ; 3_3) on the trajectory. We repeat this operation starting from 1_1 = 1_1 to nd a second point P_2 (0_1 ; 0_2 ; 3_3) on the orbit.

 $_1$, $_1^0$, $_2^0$ and $_3^0$ are chosen at random; good convergence is attained if these points belong to the m iddle zones of the variation ranges of $_1$, $_2$ and $_3$ or, at least, they are far away from the singularities on the faces of $P_3(0)$ Successive steps.

 P_1 and P_2 provide an approximation of the curve by the secant line joining them. For some small $2\ R^+$, we choose $P_3^{\ 0}=P_1+(P_2-P_1)$ as the starting value of the Newton-Raphson procedure applied to the solution of equations (35) and (36); we thus obtain the point P_3 on the curve. P_2 and P_3 lead to guesstimate by the same token another value $P_4^{\ 0}$ that produce the next point P_4 on the orbit and now the iteration is obvious. Replacing by we travelalong the opposite sense on the trajectory. The algorithm stops when one of the three variables $_1$, $_2$, $_3$ reaches its extreme value; it is applied independently on each stage, determined by the signs of $_a$ and the global trajectory is obtained by the demand for continuity.

We now describe the portrait of these orbits. Having xed $_2$ and $_3$, the corresponding kink trajectory is a non-plane curve in the interior of P_3 (0) that starts from the vacuum point D, reaches the top face BCF $_1$ O and hits the edge AF $_2$. It then goes to the edge F $_1$ F $_3$, back again to the top face, hits the edge AF $_2$ a second time, the top face a third time and ends at D: see Figure 2 and Figure 3. Varying $_2$ and $_3$ in the range of nite real numbers, other similar trajectories are obtained that hit the edges AF $_2$ and F $_1$ F $_3$ at dierent points. Given a sense of time therefore exists a two-parameter family of kink trajectories in one-to-one correspondence with the points in the interior of AF $_2$ and F $_1$ F $_3$. It should be mentioned that a whole congruence of trajectories parametrized by the interior of F $_1$ F $_3$ converges at one single point in the interior of AF $_2$ and viceversa.

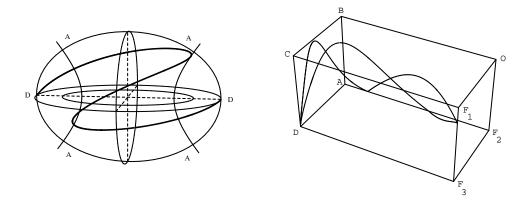


Figure 2: A generic kink drawn both in D 3 and in P $_3$ (0). Observe in D 3 that the generic kink is a heteroclinic trajectory

The translation of a generic kink trajectory to Cartesian coordinates is a delicate matter; due to the non-uniqueness of the mapping in plied by the change of coordinates special care is necessary in the analysis of the trajectory near the special conics (43)-(44) where several

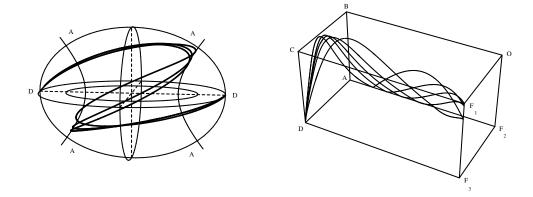


Figure 3: Several generic kinks, all of them intersecting once with the edge F_1F_3 and twice with the edge AF_2 of $P_3(0)$.

options are a priori possible. Since it requires continuity and derivability to the trajectories in the interior of the ellipsoid (40), the interior of D 3 , the behaviour of the curve is mostly xed. Two choices, the speci cation of D = \mathbf{v} as the starting point and the location of the intersection of the kink trajectory with the $\mathbf{q}_1 = 0$ plane in the quadrant characterized by $\mathbf{q}_2 > 0$, $\mathbf{q}_3 > 0$, completely x the itinerary.

There is a crossroad where the \particle" touches the $q_1 > 0$ branch of the hyperbola (44) and turns back towards the ellipse (43). There, the movement enters the $q_3 < 0$ half-space and the kink trajectory reaches the other branch, $q_1 < 0$, of the critical hyperbola after crossing the $q_2 = 0$ plane. At this stage, the particle makes its way for a third crossing of the $q_2 = 0$ plane and, nally, the journey ends at $D = \mathbf{v}^+$. This kind of kink trajectory is therefore heteroclinic: it starts and ends at different unstable points, so that:

$$Q_1^T = \frac{1}{2}^{\frac{Z}{1}} d \frac{dq_1}{d} = 1; \quad Q_2^T = Q_3^T = 0$$
 (45)

We call these topological kinks TK3 because they have three non-null components:

$$q_1() = q_1(x) \in 0; q_2() = q_2(x) \in 0; q_3() = q_3(x) \in 0$$

It should be noted that a unique, apparently non-derivable, kink trajectory in elliptic coordinates corresponds to eight derivable trajectories in Cartesian coordinates: the choices of \mathbf{v} or \mathbf{v}^+ and $\mathbf{q}_3 < 0$ or $\mathbf{q}_3 > 0$, $\mathbf{q}_2 < 0$ or $\mathbf{q}_2 > 0$ as the starting point and initial quadrant give the eight possibilities.

The energy of a three-component topological kink is the action of the trajectory times $\frac{m_p^3}{2^p \cdot \frac{1}{2}}$ and hence computable from formula (29) for the N = 3 case:

$$\frac{2^{p} \overline{2}}{m^{3}} E_{TK3} = \int_{0}^{Z} \frac{2^{2}}{p} \frac{1}{1} \frac{1}{1} + \int_{2}^{Z} \frac{2^{2}}{p} \frac{2 \cdot 2^{d} \cdot 2}{1} + \int_{2}^{Z} \frac{3 \cdot 3^{d} \cdot 3}{p} \frac{3}{1} \frac{3}{3} d \cdot 3}$$

$$= \frac{4}{3} + \frac{2^{h}}{3} \cdot 3(3 + 2^{2}) + 2(3 + 2^{2})^{\frac{1}{2}}$$
(46)

It is independent of $_2$ and $_3$ and hence the same for every kink in the TK3 fam ily.

3.3 Enveloping Kinks

There is another family of N=3 kinks living on the surface M_3 f($_1$; $_2$; $_3$)= $_1=0$ g, the unique face of $P_3(0)$ where the elliptic coordinates are not singular. In M_3 , the Hamiltonian becomes:

$$H = \frac{X^3}{{}^{0}(a)} = \frac{H_a}{{}^{0}(a)} = \frac{X^3}{{}^{0}(a)} = \frac{2A(a)}{{}^{0}(a)} = \frac{1}{a} = \frac{1}{2} = \frac{1}{2}$$

and therefore there is a two-dimensional system hidden inside the N=3 model which is H am ilton-Jacobi separable. The orbit equations,

$$C = \frac{p \frac{1}{1} \frac{2}{2} + 2}{p \frac{1}{1} \frac{2}{3} + 2} \xrightarrow{3 \text{ sign}(2)} \frac{p \frac{1}{1} \frac{2}{2} + 3}{p \frac{1}{1} \frac{2}{3} + 3} \xrightarrow{2 \text{ sign}(2)} \frac{p \frac{1}{1} \frac{2}{3} + 3}{p \frac{1}{1} \frac{3}{3} + 3}$$

$$(48)$$

are param etrized by only one real constant $_2$ (C = e $_2^{2}$ $_3^{(\frac{2}{3}-\frac{2}{2})}$).

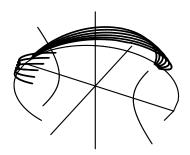


Figure 4: The NTK3 fam ily

The M athem atica plot of these solutions is shown in Figure 4. Having xed $_2$, the corresponding kink trajectory is a plane curve in M $_3$ that starts from the vacuum point D, reaches the top edge BC, goes to the umbilicus A and then back to the edge BC, to end nally in the vacuum point D. The value of $_2$ determines the points in BC where the trajectory bounces back and thus the one-parameter family of this kind of kink trajectories is in one-to-one correspondence with the points in the interior of BC.

In Cartesian coordinates the enveloping kinks are trajectories that unfold on the ellipsoid (40):

$$q_1^2 + \frac{q_2^2}{\frac{2}{3}} + \frac{q_3^2}{\frac{2}{3}} = 1$$

The starting point is either $D=v^+$ or D=v and the trajectories also end in either $D=v^+$ or D=v. A sociated with \homoclinic" trajectories, the corresponding kinks are \non-topological": $Q_1^T=Q_2^T=Q_3^T=0$. The three Cartesian components q_a dier from

zero and the appropriate name for this kind of solitary wave is a non-topological kink of three components, NTK3 for short. Every NTK3 trajectory on its way from D= \forall to D= \forall crosses the umbilicus point of the ellipsoid. Note that, again, eight trajectories in the Cartesian space R³ correspond to one trajectory in $P_3(0)$: the particle has the freedom to choose the points \forall or \forall as base points of the curve. Having xed one of them, the trajectory may develop in the half-ellipsoids determined in (45) by $q_3=0$ or $q_3=0$ and, nally, there are two travelling senses in each orbit.

Also, the energy of a three-component non-topological kink is essentially the action of the NTK3 trajectory:

$$\frac{{}_{2}P_{\overline{2}}}{{}_{m}^{3}}E_{NTK3} = \frac{{}_{2}^{Z}}{{}_{2}^{2}} \frac{{}_{2}d_{2}}{1} + \frac{{}_{2}^{Z}}{1} \frac{{}_{2}^{2}}{{}_{2}^{3}} \frac{{}_{3}d_{3}}{1} = 2 \quad {}_{2} + {}_{3} \quad \frac{{}_{2}^{Z} + {}_{3}^{Z}}{3}!$$
(49)

according to the Hamilton-Jacobi theory.

3.4 Embedded Kinks

Three-com ponent topological and non-topological kinks arise as genuine solitary waves in the N = 3 m odel. Restriction to the q_3 = 0 and/or q_2 = 0 planes shows that the N = 2 system is included twice, once in each plane, in the N = 3 m odel. Therefore, all the solitary waves of the N = 2 m odel are embedded twice as kinks of the larger N = 3 system. The embedded kinks live on the q_2 = 0 and q_3 = 0 planes, i.e. the faces of P_3 (0) where the elliptic coordinate system is singular.

I. Embedded kinks in the $q_2 = 0$ plane

Both $_2=_2^2$ and $_3=_2^2$ give $q_3=0$, see (38), and hence this coordinate plane in R 3 is the union of the two faces, $_2=_2^2$ and $_3=_2^2$, of P $_3$ (0). Therefore, in

$$M_{2_3} = {\overset{0}{\text{1}}} (_{1};_{2};_{3}) = _{3} = {\overset{0}{\overset{2}{\text{2}}}} {\overset{0}{\text{t}}} {\overset{0}{\text{(1;2;3)}}} = _{2} = {\overset{0}{\overset{2}{\overset{2}{\text{2}}}}} = M_{2_3}^{\frac{1}{3}} {\overset{1}{\text{t}}} M_{2_3}^{\frac{2}{3}}$$

we expect to nd all the kinks of the N = 2 case.

In $M_{2_3}^1$, $_3 = _2^2$, we are in the face of $P_3(0)$ such that $0 < _1 < _3^2 < _2 < _2^2$, and

$$q_1^2 = \frac{1}{\frac{2}{3}}(1 \quad _1)(1 \quad _2); \quad q_3^2 = \frac{1}{\frac{2}{3}}(\frac{2}{3} \quad _1)(\frac{2}{3} \quad \frac{2}{2})$$

The Hamiltonian also reduces to the N = 2 Hamiltonian

$$H = \frac{2A \begin{pmatrix} 1 \\ 0 \end{pmatrix} \begin{pmatrix} 2 \\ 1 \end{pmatrix}}{\begin{pmatrix} 1 \\ 2 \end{pmatrix}} \begin{pmatrix} 2 \\ 0 \end{pmatrix} \begin{pmatrix} 1 \\ 1 \end{pmatrix}} \begin{pmatrix} 2 \\ 1 \end{pmatrix} \begin{pmatrix} 2 \\ 1 \end{pmatrix} \begin{pmatrix} 2 \\ 1 \end{pmatrix} \begin{pmatrix} 2 \\ 2 \end{pmatrix} \begin{pmatrix} 1 \\ 2 \end{pmatrix} \begin{pmatrix} 2 \\ 2 \end{pmatrix} \begin{pmatrix} 2 \\ 2 \end{pmatrix} \begin{pmatrix} 2 \\ 3 \end{pmatrix}$$
$$\frac{2A \begin{pmatrix} 2 \\ 2 \end{pmatrix}}{\begin{pmatrix} 2 \\ 2 \end{pmatrix}} \begin{pmatrix} 2 \\ 2 \end{pmatrix} \begin{pmatrix} 2 \\ 2 \end{pmatrix} \begin{pmatrix} 2 \\ 3 \end{pmatrix}$$

and the Hamilton-Jacobim ethod prescribes the equation

$$e^{2 \cdot 3 \cdot \frac{2}{3} \cdot 2} = \frac{p \cdot \frac{1}{1} \cdot \frac{1}{1} \cdot 3}{p \cdot \frac{1}{1} \cdot \frac{1}{1} \cdot 1} \cdot \frac{p \cdot \frac{1}{1} \cdot \frac{1}{1} \cdot 1}{p \cdot \frac{1}{1} \cdot \frac{2}{2} \cdot 3} \cdot \frac{p \cdot \frac{1}{1} \cdot \frac{1}{2} \cdot 1}{p \cdot \frac{1}{1} \cdot \frac{2}{2} \cdot 1} \cdot \frac{1 \cdot \operatorname{sign}(\cdot_{1})}{(50)}$$

as ruling the portion of the trajectories at this face, bounded by the edges AD, AF₂, F_2F_3 and F_3D (see Figure 5).

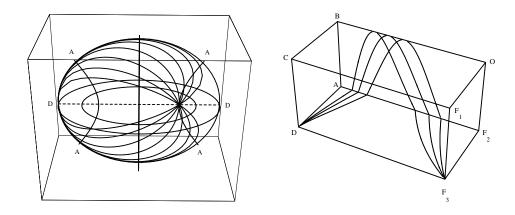


Figure 5: Several N T K 2 $_3$ kink trajectories of the N = 2 system embedded in the q_2 = 0 plane

In $M_{2_3}^2$, $_2 = \frac{2}{2}$ and the face in the boundary of $P_3(0)$ is $0 < _1 < \frac{2}{3} < \frac{2}{2} < _3 < 1$. The trajectory equations at this face are:

$$e^{2 \cdot 3 \cdot \frac{2}{3} \cdot 2} = \underbrace{\frac{p}{1} \quad \frac{1}{1} \quad \frac{3}{1}}_{p \frac{1}{1} \quad \frac{1}{3} + 3} \cdot \underbrace{\frac{p}{1} \quad \frac{1}{1} \quad \frac{1}{1}}_{p \frac{1}{1} \quad \frac{3}{3} + 1} \cdot \underbrace{\frac{1}{3}! \, sign(\ _{1})}_{sign(\ _{3})}$$

$$\underbrace{\frac{p}{1} \quad \frac{3}{3} \quad \frac{3}{3}}_{p \frac{1}{1} \quad \frac{3}{3} + 3} \cdot \underbrace{\frac{p}{1} \quad \frac{3}{3} + 1}_{p \frac{1}{1} \quad \frac{3}{3} \quad 1} \cdot \underbrace{\frac{1}{3}! \, sign(\ _{3})}_{(51)}$$

and the boundary is form ed by the edges AF_2 , F_2O , OB and BA.

For nite values of $_2$, som e kink trajectories given by (50)–(51) are depicted in Figure 5. It may be observed that the trajectory starts at D and then runs through the face M $_2^1$ until the edge A $_2^2$. From this point, the particle enters the M $_2^2$ face (here the path is not derivable), reaches the BO edge and comes back to the A $_2^2$ face. This is the second point of non-di erentiability re-entering the trajectory the M $_2^1$ face. All the trajectories then meet at the vertex $_3^2$ and come back in a symmetric way to end in the D point. In Cartesian coordinates, these kink trajectories start and end in either D = $_3^2$ or D = $_3^2$, $_3^2$ do not leave the $_3^2$ plane, and cross either the focus ($_3^2$ point in Cartesian therefore call them NTK2 $_3^2$ because they are two-component non-topological kinks, merely the family of NTK2 kinks of the N = 2 m odel, embedded this way within the manifold of

kinks of the N = 3 system . There are four trajectories of this kind inside the ellipsoid (45) in R 3 per trajectory in the boundary of P $_3$ (0): there is freedom to choose v^+ or v^- and the sense of travel in each orbit. The NTK 2 $_3$ kinks are xed points of the Z $_2$ sub-group of G = Z $_2$ 3 generated by q_2 ! q_2 , that, however, does not leave invariant the TK 3 and the NTK 3 trajectories . The energy of these solutions is:

$$\frac{2^{p} \overline{2}}{m^{3}} E_{NTK2_{3}} = \frac{\sum_{3}^{2} \frac{1}{2} \frac{1}{1} \frac{1}{1}}{\sum_{1}^{2} \frac{1}{1}} + \frac{\sum_{2}^{2} \frac{2}{2} \frac{2 \cdot 2 \cdot 2}{2}}{\sum_{1}^{2} \frac{1}{1} \frac{2}{2}} + \frac{\sum_{1}^{2} \frac{2}{3} \frac{3 \cdot 3}{2}}{\sum_{1}^{2} \frac{1}{1} \frac{3}{3}}$$

$$= \frac{4}{3} + 2 \cdot 3 \cdot 1 \cdot \frac{3}{3} \tag{52}$$

There is a lim iting case to this fam ily of kinks: a trajectory along the DA and AB edges and back to D through the same way. Elliptic coordinates are even more singular on the edges, but the dynamical system reduces to a one-dimensional Hamiltonian system which can be integrated analytically. We have a two-step trajectory:

At the DA edge, $_1$ = 0 and $_3$ = $_2^2$, the canonical equations (after use of the rst integral) reduce to:

$$\frac{d_{2}}{d} = 2(_{2} \quad _{3}^{2}) \frac{q}{1}$$
 (53)

with the solution

$$_{1}^{TK2}$$
 () = 0; $_{2}^{TK2}$ () = 1 $_{3}^{2}$ tanh²($_{3}$); $_{3}^{TK2}$ () = $_{2}^{2}$ (54)

for 2 (1; $\frac{1}{3}$ arctanh $\frac{2}{3}$] t [$\frac{1}{3}$ arctanh $\frac{2}{3}$;1). The second step occurs on the AB edge, where, again, the canonical equations reduce to a single di erential equation: if $_1 = 0$ and $_2 = \frac{2}{3}$,

$$\frac{d_{3}}{d} = 2(_{3} \quad _{3}^{2})^{1} \quad \frac{1}{1} \quad _{3}$$
 (55)

has the solution

$$_{1}^{TK2_{3}}() = 0;$$
 $_{2}^{TK2_{3}}() = _{2}^{2};$ $_{3}^{TK2_{3}}() = 1$ $_{3}^{2} \tanh^{2}(_{3})$ (56)

for $2^{\frac{1}{3}}$ arctanh $\frac{2}{3}$; $\frac{1}{3}$ arctanh $\frac{2}{3}$. The corresponding kinks in Cartesian coordinates are TK2 $_3$ and TK2 $_3$, the four two-component topological kinks of the N = 2 m odel:

Thus, the enveloping kinks of the $N=2\,m$ odelare also embedded in the N=3 system. The energy for these solutions and their anti-kinks is:

$$\frac{{}_{2}^{p}\overline{2}}{{}_{m}^{3}}E_{TK2_{3}} = \frac{{}_{2}^{2}}{{}_{3}^{2}} \frac{{}_{2}^{2}}{{}_{1}^{2}} + \frac{{}_{2}^{2}}{{}_{2}^{2}} \frac{{}_{2}^{3}}{{}_{1}^{3}} = 2_{3} 1 \frac{{}_{3}^{2}}{3}$$
(58)

In the $q_2 = 0$ plane there is still one trajectory that is even m ore singular: it is a three step path running on the edges DF₃, F₃F₂, F₂O and back to D through the same way. The canonical equations and its solutions in the three steps are:

1.
$$_{2} = _{3}^{2}$$
 and $_{3} = _{2}^{2}$.

$$\frac{d_{1}}{d} = _{2}^{1} \frac{q_{1}}{1}; \quad 2 (1 ; arctanh_{3}]t [arctanh_{3};1)$$

$$_{1}^{TK1}() = 1 tanh^{2} ; _{2}^{TK1}() = _{3}^{2}; _{3}^{TK1}() = _{2}^{2}$$
2. $_{1} = _{3}^{2}$ and $_{3} = _{2}^{2}$.

$$\frac{d_{2}}{d} = _{2}^{2} \frac{q_{1}}{1}; \quad 2 [arctanh_{3}; arctanh_{2}]t [arctanh_{2}; arctanh_{3}]$$

$$_{1}^{TK1}() = _{3}^{2}; _{2}^{TK1}() = 1 tanh^{2}; _{3}^{TK1}() = _{2}^{2}$$
3. $_{1} = _{3}^{2}$ and $_{2} = _{2}^{2}$.

$$\frac{d_{3}}{d} = _{2}^{3}; _{3}^{TK1}() = _{2}^{2}; _{3}^{TK1}() = _{2}^{2}; _{3}^{TK1}() = _{1}^{2}; arctanh_{2}]$$

Only one Cartesian component is dierent from zero:

and hence the one-component topological kink of the $N=1\,m$ odel is embedded rst in the manifold of kinks of the $N=2\,m$ odel, and then in the N=3 system. There are two kinks of this kind in Cartesian coordinates which are mapped in a unique trajectory in the boundary of $P_3(0)$. The TK1 trajectories are xed points of the Z^2 sub-group of G generated by $Q_2!$ Q_2 and $Q_3!$ Q_3 . The energy is

$$\frac{2^{p} \overline{2}}{m^{3}} E_{TK1} = \int_{0}^{Z} \frac{1}{p} \frac{1}{1} \frac{1}{1} + \int_{2}^{Z} \frac{2}{p} \frac{2}{1} \frac{2}{2} + \int_{2}^{Z} \frac{1}{p} \frac{3d \ 3}{1} \frac{3}{3} = \frac{4}{3}$$
 (60)

Embedded K inks in the $q_3 = 0$ plane

The DF₃, F₃F₂ and F₂O edges form the intersection of the $q_2 = 0$ and $q_3 = 0$ planes. Therefore, the TK1 kinks also live in the $q_3 = 0$ plane. There is another maximally singular trajectory living on the \edge" in the $q_3 = 0$ plane:

At the DC edge, $_2=\frac{2}{3}$ and $_1=0$, the canonical equations for the nite action trajectories are:

$$\frac{d_{3}}{d} = 2(_{3} \quad _{2}^{2})^{1} \quad _{3}$$
 (61)

The path

$$_{1}^{\text{TK 2}}$$
 () = 0; $_{2}^{\text{TK 2}}$ () = $_{3}^{2}$; $_{3}^{\text{TK 2}}$ () = 1 $_{2}^{2}$ tanh² ($_{2}$) (62)

solves (61) and runs when goes from 1 to +1 from D to D passing through the vertex C at = 0. In C artesian coordinates we recover the four two-component topological kinks of the N = 2 m odel, now embedded in the $q_B = 0$ plane:

These are heteroclinic trajectories that produce the TK2 $_2$ and TK2 $_2$ topological kinks. The energy is:

$$\frac{{}_{2}P_{\overline{2}}}{{}_{m}^{3}}E_{TK2_{2}} = {}_{2}^{Z_{1}} \frac{{}_{3}d_{3}}{\overline{1}_{3}} = 2_{2} 1 \frac{{}_{2}^{2}!}{\overline{3}}$$
(64)

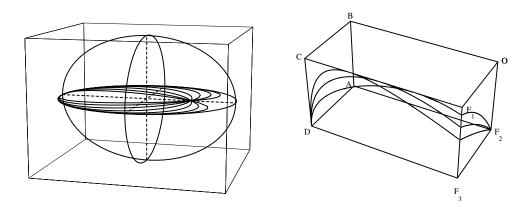


Figure 6: NTK2 $_2$ kink trajectories of the N = 2 system embedded in the q_3 = 0 plane

Of course, the full manifold of kinks of the $N=2\,m$ odel is embedded in the $q_3=0$ plane: the set of kinks of the N=3 system is completed by the two-component non-topological kinks living in the $q_3=0$ plane, see Figure 6. The $q_3=0$ plane is mapped to the union of two faces in the boundary of $P_3(0)$:

$$M_{2_{2}} = {\atop }^{n} (_{1};_{2};_{3}) = _{1} = {\atop }^{2} {\atop }^{0} t {\atop }^{n} (_{1};_{2};_{3}) = _{2} = {\atop }^{2} {\atop }^{0} = M_{2_{2}}^{1} t M_{2_{2}}^{2}$$

1. In $M_{2_2}^1$, $_1 = \begin{array}{cc} 2 \\ 3 \end{array}$ im plies $q_3 = 0$. Therefore:

$$q_1^2 = \frac{1}{\frac{2}{2}}(1 \quad 2)(1 \quad 3)$$
 $q_2^2 = \frac{1}{\frac{2}{2}}(\frac{2}{2} \quad 2)(\frac{2}{2} \quad 3)$

is a well de ned change of coordinates in the range $\frac{2}{3} < \frac{2}{2} < \frac{2}{2} < \frac{2}{3} < 1$. In this region, the interior of the ellipse (43), the trajectories providing kinks are given by the

equations:

$$e^{2 \cdot 2 \cdot \frac{2}{2} \cdot 2} = \frac{p \cdot \frac{1}{1 \cdot 2} \cdot 2}{p \cdot \frac{1}{1 \cdot 2} \cdot 2} = \frac{p \cdot \frac{1}{1 \cdot 2} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} = \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{(65)}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} = \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1 \cdot 3} \cdot 1}{p \cdot \frac{1}{1 \cdot 3} \cdot 2} \cdot \frac{p \cdot \frac{1$$

2. In $M_{2_2}^2$, $_2=$ $_3^2$ also in plies $q_3=$ 0. In the range 0 < $_1<$ $_3^2<$ $_2^2<$ $_3<$ 1 the change of coordinates is de ned as

$$q_{1}^{2} = \frac{1}{\frac{2}{2}}(1 \quad 1)(1 \quad 3)$$

$$q_{1}^{2} = \frac{1}{\frac{2}{2}}(\frac{2}{2} \quad 1)(\frac{2}{2} \quad 3)$$

The kink trajectories satisfy the equations:

$$e^{2 \cdot 2 \cdot \frac{2}{2} \cdot 2} = \frac{p \cdot \frac{1}{1} \cdot \frac{1}{1} \cdot 2}{p \cdot \frac{1}{1} \cdot \frac{1}{1} \cdot 1} \cdot \frac{p \cdot \frac{1}{1} \cdot \frac{1}{1} \cdot 1}{p \cdot \frac{1}{1} \cdot \frac{3}{3} \cdot 2} \cdot \frac{p \cdot \frac{1}{1} \cdot \frac{1}{3} \cdot 1}{p \cdot \frac{1}{1} \cdot \frac{3}{3} \cdot 1} \cdot \frac{p \cdot \frac{1}{1} \cdot \frac{3}{3} \cdot 1}{p \cdot \frac{1}{1} \cdot \frac{3}{3} \cdot 1} \cdot \frac{1}{2} \cdot \frac{1}{3} \cdot \frac{1}{3$$

The features of this kind of kinks are identical to the characteristics of the two-com ponent non-topological kinks that exist in the $q_2=0$ plane. The only dierence is that they have support in the faces M $_2^1$ and M $_2^2$ instead of M $_2^1$ and M $_2^2$ and we therefore call them NTK2 $_2$. They meet at the vertex F $_2$, and therefore at the foci ($q_1=2$;0;0) in R $_2^3$; see Figure 6. The energy is:

$$\frac{{}_{2}^{p}\overline{2}}{{}_{m}^{3}}E_{NTK2_{2}} = 2 \int_{0}^{"} \frac{{}_{2}^{2}}{{}_{2}^{3}} \frac{{}_{1}^{d}}{{}_{1}^{1}} + 2 \int_{2}^{2} \frac{{}_{2}^{2}}{{}_{3}^{2}} \frac{{}_{2}^{d}}{{}_{2}^{2}} + 2 \int_{2}^{2} \frac{{}_{3}^{d}}{{}_{3}^{3}} \frac{{}_{3}^{d}}{{}_{3}^{3}} =$$

$$= \frac{4}{3} + 2 \int_{2}^{2} \frac{{}_{3}^{2}}{{}_{3}^{2}} \frac{{}_{2}^{d}}{{}_{3}^{2}} \frac{{}_{2}^{d}}{{}_{2}^{2}} + 2 \int_{2}^{2} \frac{{}_{3}^{d}}{{}_{3}^{3}} \frac{{}_{3}^{d}}{{}_{3}^{2}} =$$

$$(67)$$

In sum: the manifold of kinks of the N = 2 m odel is embedded twice in the N = 3 system, once in the q_2 = 0 plane and other in the q_3 = 0 plane. They are sewn togehter by the common TK1, embedded from the N = 1 m odel. The embedded kinks llthe gaps left by the TK3 families of kinks in the interior of the ellipsoid (45) and also develop through the curves left by the NTK3 families on the boundary of D³. D³ is thus a \totally" geodesic manifold with respect to the separatrices between bounded and unbounded motion in the N = 3 dynamical system. The NTK3 family form the envelop of the separatrices and the NTK3 kinks are them selves separatrices.

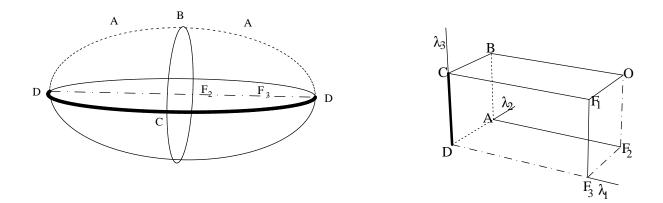


Figure 7: Plot of the \singular" topological kinks TK2 $_2$ (solid line), TK2 $_3$ (broken line) and TK1 (dash-dotted line) in both Cartesian and Elliptic coordinates.

4 Further Comments

We now infer the general structure of the kink manifold of the linear 0 (N)-sigm a model from the pattern shown by the 0 (2)-and 0 (3)-sigm a models. We can safely state that all the kink trajectories live in the sub-manifold $D^N = R^N$ determined by the inequality:

$$q_1^2 + \frac{q_2^2}{\frac{2}{2}} + \dots + \frac{q_2^2}{\frac{2}{N}}$$
 1

There are three categories:

1. Generic Kinks

A . There exists a fam ily of generic kinks param etrized by N $\,$ 1 real constants that live in the interior of D $^{\rm N}$. The intersection loci of the generic kinks are the singular quadrics:

$$\frac{q_{1}^{2}}{\frac{2}{N}} + \frac{q_{2}^{2}}{\frac{2}{N}} + \frac{q_$$

B. Generic kinks are non-topological-hence NTKN-ifN even, and topological-hence TKN-ifN is odd.

2. Enveloping kinks.

A . The restriction of the dynam ical system to the boundary @D $^{\rm N}$ of D $^{\rm N}$, the hyperellipsoid:

$$q_1^2 + \frac{q_2^2}{\frac{2}{2}} + \dots + \frac{q_2^2}{\frac{2}{2}} = 1$$

provides a fam ily of enveloping kinks param etrized by N = 2 real constants. Recalling that in elliptic coordinates @D $^{\rm N}$ is characterized by the equation $_{1} = 0$, the intersection loci of this congruence are the um bilical sub-m anifolds:

$$_{1} = 0$$
; $_{a} = {}^{2}_{N} {}_{a+1} = {}_{a+1}$; $8a = 2;3;...;N$

of dim ension N 3 of the hyper-ellipsoid @D $^{\rm N}$.

B. Enveloping kinks are topological, hence TKN, if N is even, and non-topological, hence NTKN, if N is odd.

3. Embedded Kinks.

On the N $1\,\mathrm{R^{\,N^{\,-1}}}$ sub-m anifolds determ ined by the conditions $q_a=0$, if a=2 or 3 or ::: or N , the dynam ical system reduces to the mechanical system that arises in the linear O (N 1)-sigm a model. Thus, the kink manifold of the N 1 case is included N 1 times in the O (N)-model, lling the holes left in the interior of D $^{\mathrm{N}}$ by the generic kinks, and also covering in $0\,\mathrm{C}$ the sub-spaces which are not covereded by the enveloping kinks. Each N $1\,\mathrm{kink}$ sub-manifold is not, however, included (N 1) (N 2) times in the O (N)-model because the R $^{\mathrm{N}}$ 2 sub-spaces are intersections of the N $^{\mathrm{N}}$ 2 , dened above. The N $^{\mathrm{N}}$ 1 kink manifolds are not separated but sewn together through the N $^{\mathrm{N}}$ 1 kink sub-manifolds. This is a iterative process in such a way that the kink manifold of the O (N $^{\mathrm{N}}$)-sigm a model is included $^{\mathrm{N}}$ $^{\mathrm{N}}$ $^{\mathrm{N}}$ times in the kink manifold of the O (N $^{\mathrm{N}}$) system .

O ne can ask what happens if a continuous sub-group O (r) of O (N) survives as symmetry group of the system . This happens if the deformation is chosen in such a way that $0 < \frac{2}{2} = \frac{2}{3} = \frac{2}{r} < \frac{2}{r+1} < :::< \frac{2}{N} < 1$. In this case we obtain a sub-manifold of kinks from O (r) rotations around the q_1 axis of the kink manifold of the N = 2 system that lives in the q_1 : q_2 plane. The remaining kinks correspond to the solitary waves of the N = r 1 system dened in the orthogonal R N r+1 sub-space. Also the deformations where $1 < \frac{2}{r+1} < \frac{2}{N}$ are easy to understand. Finite action trajectories spread out in the domain in R N bounded by the hyper-hyperboloid:

$$q_1^2 + \frac{q_2^2}{\frac{2}{2}} + \frac{q_1^2}{\frac{2}{2}} + \frac{q_{r+1}^2}{\frac{1}{2}} = 1$$
:

The kink manifold of this system is the kink manifold of the N = r model de ned in the sub-space R $^{\rm r}$ R $^{\rm N}$ such that $q_{r+1} = \frac{1}{N} \neq 0$.

Finally we consider a mild deformation of our model by introducing asymmetries in the non-harmonic terms of the potential energy and also adapting the quadratic terms in a suitable manner:

The new non-dimensional constants $"_a$ and $_a$ are defined in terms of the old $_a$'s through:

$$1 + \mathbf{a} = \frac{a(a+1)}{2}$$
; $a = \frac{3(2+a)}{2}$; $a = 2;3;...;N$

Am ong the kinks of the deform ed linear O (N)-sigm a m odel only the following survive as solitary wave solutions of this perturbed system:

(a) The TK 1.

$$_1 = \tanh x$$
; $_2 = :::= N = 0$

(b) All the TK2 kinks. On the ellipse,

$${\begin{array}{c} 2 \\ 1 \end{array}} + \frac{1 + {\begin{array}{c} 1 \\ 2 \\ 3 \end{array}}}{1 + {\begin{array}{c} 2 \\ 3 \\ 3 \end{array}}} = 1$$

the TK2 $_{\rm a}$ and TK2 $_{\rm a}$ con gurations,

$$a = \tanh_{a} x ; \quad a = \frac{1 + \frac{2}{a}}{1 + \frac{1}{a}} \text{ sech } ax ; \quad b = 0 ; 8b \in a; b \in 1$$

are solutions of the eld equations. The amazing fact is that in this deformation of the O (N)-linear sigma model, discussed by Bazeia et al. if N=2 [12], the energy of all these kinks is the same:

$$E_{TK1} = E_{TK2_2} = ::= E_{TK2_N} = \frac{4}{3^2} \frac{m^3}{2}$$

On one hand, we have a deform ation of the linear O (N)-sigm a model that exhibits a complex variety of kinks; on the other hand, another deform ation of the O (N)-model rejects almost every kink but the simplest ones as solutions, and all of the surviving kinks are degenerated in energy.

5 Outlook

The developm ents disclosed in this paper suggest a general strategy in the search for kinks in two space-timedimensional eld theories. When the elds have N components assembled in a vector representation of the O (N) group, we focus on systems with symmetry breaking to a discrete sub-group of O (N) which has more than one element. If the dynamical system that determines the localized static solutions is completely integrable, all the solitary waves can be found, at least in principle. Particularly interesting is the situation where the N 1 invariants in involution with the mechanical energy act non-trivially on the manifold of localized solutions and the orbit is a continuous space. One can then perturb such a system, loosing in the perturbation many of the solitary wave solutions: only few of the localized static solutions survive as kinks of the perturbed (m ore realistic) model.

We nally list several interesting questions that will be postponed for future research:

study of the structure of the kink m anifold of the deform ed linear O (N)-sigm a m odel as a m oduli space seem s to be worthwhile.

A detailed analysis of the sum rules between the energies of the dierent kinds of kinks is necessary to x the structure mentioned above.

A treatm ent a la Bogom olny is also possible. This allows for a supersymmetric extension of the model in such a way that the kinks become BPS states.

The dicult problem remains of determining the stability of the dierent kinds of kinks. Application of the Morse index theorem helps in noting the stability properties which in turn provide information about the quantization of these topological defects.

6 Acknowledgements

The authors are grateful to Askold Perelom ov for teaching them the magic of the elliptic Jacobi coordinates and their relationship to dynamical problems on ellipsoids.

A ppendix: Elliptic coordinates

G iven any set of N real positive numbers such that $0 < r_1 < r_2 < :::< r_N$, let us consider the equation:

$$\frac{x^{N}}{x_{a}} = \frac{q_{a}^{2}}{r_{a}} = 1 \tag{68}$$

The left-hand member Q (q) = $\frac{x^N}{a=1} \frac{q_a^2}{r_a}$ of this equation can be understood either as a function of R N , for xed 2 C, or as a function of the complex variable, for xed q 2 R N . From (68) one immediately deduces:

1 Q
$$(q) = 1 + {x^{N} \over a=1} \frac{q_{a}^{2}}{r_{a}} = {x^{N} \over x^{N} \over x^{N}} = 0$$
 (69)

Therefore, the N roots $_a$ of the polynom ial in the numerator of 1 $\,$ Q (q), a rational function of ,are the roots of equation (68). The roots $_a$ are also real numbers and 1 Q (q) is a rational function such that $_1 < r_1 < _2 < ::: < r_{\text{N}} \ _1 < _{\text{N}} < r_{\text{N}}$, see Figure 8. To prove this point one needs to study 1 $\,$ Q (q) along the $\,$ -real axis, near the poles $\,$ = $\,$ r_a, using Bolzano's theorem .

De nition. The elliptic coordinates of the point q $(q_1; :::; q_N)$ 2 R N are the roots $^\sim_E$ $(q_1; :::; q_N)$ 2 P N (1) of Q (q) = 1.

 P^N (1) R^N is the open sub-space of R^N given by: $1 < {}_1 < r_1, r_1 < {}_2 < r_2, \ldots, r_{N-1} < {}_N < r_N$. The solution of (68) for $= {}_1$ constant 2 (1; r_1) is, geometrically, a quadric surface, a hyperellipsoid, when q varies in R^N . The family of quadrics obtained by taking $= {}_a$ constant 2 (r_{a-1} ; r_a), a 2, corresponds to a family of hyper-hyperboloids of every possible signature in R^N , if Q (q) is considered as a function of q.

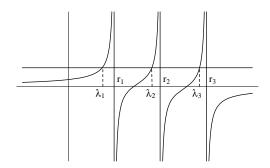


Figure 8: Plot of the function $y = \frac{x^3}{r_a} \frac{q_a^2}{r_a}$ and the function y = 1, see (68), for xed values of q_a and r_a .

It is convenient to denote the products in (69) as:

$$() = {\overset{{}_{\scriptstyle Y}}{\overset{}_{\scriptstyle A}}} (\qquad {}_{a}) ; \qquad A () = {\overset{{}_{\scriptstyle Y}}{\overset{}_{\scriptstyle A}}} (\qquad {}_{r_{a}})$$

so that

$$1 + \sum_{a=1}^{X^{N}} \frac{q_{a}^{2}}{r_{a}} = \frac{q_{a}^{2}}{Y^{N}} \frac{r_{a}}{r_{a}} = \frac{r_{a}}{Y^{N}} \frac{r_{a}}{r_{a}} \frac{r_{a}}{r_{a}} = \frac{r_{a}}{r_{a}} \frac{r_{a}}{r$$

An explicit formula for de ning q_a^2 as a function of the 's,8a, is obtained by applying the residue theorem to both mem bers of the equation (70):

Res (1 Q (q))
$$j_{=r_a} = q_a^2$$

Res $\frac{()}{A()} = \frac{(r_a)}{A^0(r_a)}; A^0(r_a) = \frac{dA()}{d}$

= ro

Therefore:

$$q_{a}^{2} = \frac{(r_{a})}{A^{0}(r_{a})} = \frac{\frac{\dot{Y}^{N}}{(r_{a} + b)}}{\frac{\dot{Y}^{N}}{\dot{Y}^{N}}}$$

$$(r_{a} + r_{b})$$

$$(72)$$

$$b = 1,56 a$$

and we see that the transform ation $(q_1; ::::; q_N)$! (1; ::::; N) is 2^N to 1.

Inverting (72) to express a as a function of the q's, 8a, requires one to solve an algebraic equation in a with powers up to a. This is easy for a = 2, possible, but very dicult for a = 3;4 using Cardano's form ulas, and in possible if a 5. For this reason, another derivation of (72) is useful, which in passing allows one to show identities between Cartesian

and elliptic coordinates that make practical computations possible. To do this, notice that (70) implies:

Setting = r_c in (73) one im m ediately derives (72). M ore important, expanding the two m em bers of (73) in a power series in , we obtain:

Equalizing the coe cients of the term swith the same power of in the last equation we have N non trivial identities. We shall use the equalities between the coe cients of $^{N-1}$ and $^{N-2}$:

A nother important tool using elliptic coordinates is the Jacobi lem m a: Lem m a. The expression

$$X^N$$
 $\begin{array}{c} & & s \\ & & a \end{array}$ $a=1$ $(a 1)(a 2):::(a):::(a N)$

where ($_1$;:::; $_N$) are N real num bers such that $_1$ < $_2$ < ::: < $_N$ is equal to 0 if s N 2 and 1 for s = N 1. Proof:

Consider the function

$$f_{s}(z) = \frac{z^{s}}{y^{N}}$$

$$(z a)$$

$$= 1$$

which has N poles in the complex plane at z=a, 8a, and another pole at in nity. If is a closed curve which is the boundary of a region D of the complex plane containing all the nite poles of $f_s(z)$, the residue theorem tells us that:

$$\frac{1}{2 \text{ i}}^{\text{I}} f_s(z)dz = \sum_{a=1}^{X^{\text{N}}} Res(f_s)(a) = Res(f_s)(1) = \frac{1}{2 \text{ i}}^{\text{I}} f_s(z)dz$$

Also,

$$\begin{array}{c} x^{N} & \text{Res}(f_{s})(a) = \\ x^{n} & & \\ x^{n}$$

and the lem m a is proved.

Applying this result to the choice a = a, 8a = 1; ...; N , we obtain new identities

$$\frac{x^{N}}{a=1} - \frac{s}{0(a)} = 0;8s < N \qquad 1$$
 (76)

$$\frac{x^{N}}{a=1} - \frac{\frac{N}{a}}{\binom{a}{0}} = 1$$
(77)

because $^{0}(_{a}) = (_{a} _{1})(_{a} _{2}) ::: (_{a} _{N})$. A lternatively if we take $_{a} = r_{a}$, the lemma in plies that:

$$\sum_{a=1}^{x^{N}} \frac{r_{a}^{s}}{A^{0}(r_{a})} = 0;8s < N \qquad 1; \qquad \sum_{a=1}^{x^{N}} \frac{r_{a}^{N-1}}{A^{0}(r_{a})} = 1$$
 (78)

An important identity obtained from the lemma is:

$$\frac{x^{N}}{a=1} \frac{q_{a}^{2}}{(r_{a} b)(r_{a} c)} = 0; 8b; c$$
(79)

The normal vectors to the family of quadrics (68)

Q (q) =
$$\frac{x^N}{r_a} \frac{q_a^2}{r_a} = 1$$

at the point q $(q_1; :::; q_N)$, are

$$n()$$
 $(n_1()); :::; n_N()) = \frac{q_1}{r_1}; :::; \frac{q_N}{r_N}$

Observe that (79) im plies

Therefore, all the quadrics are orthogonal with each other and the elliptic coordinates form an orthogonal system. The standard Euclidean metric in Cartesian coordinates can be expressed in elliptic coordinates in the form:

$$ds^2 = \int_{a=1}^{X^N} dq_a^2 = \int_{a=1}^{X^N} \int_{b=1}^{X^N} g_{ab} (\tilde{b}_E) d_a d_b$$

Derivation of the two members of equation (72) leads to:

$$\frac{2dq_a}{q_a} = (1)^N \int_{b=1}^{x^N} \frac{d_b}{r_a}$$

and, using the Jacobi Lem ma,

$$4dq_{a}^{2} = q_{a}^{2} \sum_{b=1}^{X^{N}} \frac{d_{b}^{2}}{r_{a}}$$

Finally, we have:

$$g_{aa} = \frac{1}{4} \frac{{}^{0}(a)}{A(a)} = \frac{1}{4} \frac{{}^{b6}a = 1}{{}^{W}}; g_{ab} = 0;8a \in b$$

$$(80)$$

The kinetic energy of a natural dynamical system in elliptic coordinates is;

$$T = \frac{1}{2} \sum_{a=1}^{X^{N}} q_{a}^{2} = \frac{1}{2} \sum_{a=1}^{X^{N}} q_{aa} - \frac{2}{a}$$
 (81)

In term s of the canonical m om entum $_{a}=\frac{\varrho_{T}}{\varrho_{-a}}$, T reads:

$$T = \sum_{a=1}^{X^{N}} a^{-a} \qquad T = \frac{1}{2} \sum_{a=1}^{X^{N}} \frac{1}{g_{aa}} \sum_{a=1}^{2} 2 \sum_{a=1}^{X^{N}} \frac{A(a)}{g(a)} \sum_{a=1}^{2} (82)$$

R eferences

- [1] S.Coleman, \Classical Lumps and their Quantum Descendants", in \New Phenomena in Subnuclear Physics", A.Zichichi editor, Erice Lecture Notes, Part A, Plenum Press, New York, 1976.
- [2] R.Rajaram an, \Solitons and Instantons", North-Holland, Am sterdam, 1982.
- [3] C.Montonen, Nucl. Phys. <u>B112</u> (1976), 349. S. Sarker, S. Trullinger and R. Bishop, Phys. Lett. <u>A59</u> (1976), 255.
- [4] E.M agyari and H. Thom as, Phys. Lett. <u>A 100</u> (1984), 11.
- [5] H. Ito, Phys. Lett. <u>A 112</u> (1985), 119.
- [6] A. Alonso Izquierdo, M. A. Gonzalez Leon and J. Mateos Guilarte, J. Phys. A: Math. Gen. 31 (1998), 209.
- [7] J. Liouville, J. M ath. Phys. Appl. 11 (1849), 345.
 - A. Perelomov, \Integrable Systems of Classical Mechanics and Lie Algebras", Birkhauser, Boston MA., 1990.
- [8] A. Alonso Izquierdo, \One dimensional Solitons in Complex Scalar Field Theory", Tesina de Grado, Salam anca University, 1995.
- [9] S.Colem an, Commun. Math. Phys. 31 (1973), 259
- [10] A. Zam olodchikov, Adv. Stud. Pure Math. 19 (1989), 642.
- [11] A. Alonso Izquierdo, M. A. Gonzalez Leon and J. Mateos Guilarte, \Kinks out of Geodesics. Topological walls in the Linear Sigma Model". A nales de F. sica, Monograf as 5, (1998) 15, Ed. CIEMAT, Madrid.
- [12] D. Bazeia, J. R. S. Nascim ento, R. F. Ribeiro and D. Toledo, J. Phys. A: Math. Gen. 30 (1997) 8157.
- [13] R. Jackiw and R. Schrie er, Nucl. Phys. B 190 (1981), 253.
- [14] A.Vilenkin and E.P.S. Shellard, Cosmic Strings and other Topological Defects, Cambridge University Press, 1994.
- [15] G.D valiand M. Shifm an, Nucl. Phys. <u>B504</u> (1997), 127.
- [16] M.Gell-Mann and M.Levy, Il Nuovo Cim. 16 (1960), 705.
- [17] M. Veltman, \Re ections on the Higgs systems", CERN Yellow report 97-05, 1997.
- [18] D. Bazeia and J. Morris, Phys. Rev. D 54 (1996), 5217.
- [19] J.Morris, Phys. Rev. <u>D 51</u> (1995), 697.

- [20] R.Gamier, Ren.Circ.Mat.Palemo 43 (1919), 155.
- [21] H.Grosse, Acta Phys. Austriaca <u>52</u> (1980), 89.
- [22] D.Olive, N. Turok and J. Underwood, Nucl. Phys <u>B401</u>(1993), 663.